## Molecular Electronic Structure

## Quantum Supplementary Hilary Term 2020

## Molecular wavefunctions

1. Introduction: Schrödinger equation, Born-Oppenheimer approximation, wavefunctions.
2. The linear combination of atomic orbitals (LCAO) approach to molecular orbitals
3. Pauli principle, Slater determinants

## Formulation of the Hamiltonian for $\mathbf{H}_{\mathbf{2}}{ }^{\boldsymbol{+}}$

4. Matrix formulation and the secular determinants for a 1-electron system $\left(\mathrm{H}_{2}{ }^{+}\right)$
5. Formulation of $\mathrm{H}_{2}{ }^{+}$in the LCAO approach

Basis sets
6. Slaters vs Gaussians
7. Hierachy of basis sets

Formulation of the Hamiltonian for $\mathbf{H}_{2}$.
8. Coulomb integrals
9. Triplet state of $\mathrm{H}_{2}$ and the exchange integral.
10. Expansion of the Coulomb and exchange integrals in an atomic basis.

## Self-consistent fields and Hartree-Fock theory

11. Hartree-Fock and Roothaan equations
12. An application:the origin of effective nuclear charge

Beyond the HF approximation
13. Configuration interaction
14. Density functional theory

Semi-empirical methods
15. Hückel and Extended Hückel theory

## Bibliography

1. Richards and Cooper, Ab Initio Molecular Orbital Calculations for Chemists.
2. Atkins and Friedman, MQM, $5^{\text {th }}$ ed
3. Lowe, Quantum Chemistry, $2^{\text {nd }}$ ed
4. McQuarrie, Quantum Chemistry, $2^{\text {nd }}$ ed

## Some basics:

In studies of molecular electronic structure, we seek solutions of the time-independent Schrödinger equation,

$$
-\frac{\hbar}{2 m} \nabla^{2} \varphi+V \varphi=E \varphi
$$

where $\boldsymbol{m}$ is the mass of an electron, $\boldsymbol{V}$ is the potential and $\nabla^{2}$ is the Laplacian operator

$$
\nabla^{2}=\frac{\partial^{2}}{\partial x^{2}}+\frac{\partial^{2}}{\partial y^{2}}+\frac{\partial^{2}}{\partial z^{2}}
$$

The wavefunction, $\boldsymbol{\varphi}$, must be single-valued, finite and continuous.

## Born-Oppenheimer approximation

The above equation is derived from the complete (electronic + nuclear) Schrödinger equation by regarding the nuclei as being fixed in space, and so the total wavefunction factorises as

$$
\Xi(R, r)=\psi(r) \Omega(R)
$$

The electronic Hamiltonian, $\hat{\boldsymbol{H}}$, for a system with $n$ electrons and $N$ nuclei is then given by

$$
\widehat{H}=-\frac{\hbar^{2}}{2 m} \sum_{i}^{n} \nabla_{i}^{2}-\frac{e^{2}}{4 \pi \varepsilon_{0}} \sum_{i}^{n} \sum_{a}^{N} \frac{Z_{a}}{r_{i a}}+\frac{e^{2}}{4 \pi \varepsilon_{0}} \sum_{i}^{n} \sum_{j>i}^{n} \frac{1}{r_{i j}}+\frac{e^{2}}{4 \pi \varepsilon_{0}} \sum_{a}^{N} \sum_{b>a}^{N} \frac{1}{R_{a b}}=E \varphi
$$

The final term is independent of the position of the electrons, and adds a constant contribution to the energy (at fixed distance) This defines the potential energy curve: the electronic energy which depends parametrically on the nuclear coordinates.


## Linear Combination of Atomic Orbitals (LCAO) approximation

Molecular orbitals are typically expanded as linear combinations of their atomic components:
$\varphi=\sum_{a} c_{a} \chi_{a}$

## Slater determinants (see Valence Lecture 3, pp 12)

The Pauli principle: The total wavefunction must be antisymmetric under the exchange of identical particles.

Example: For He, the 2-electron wavefunction $\Phi=\mathbf{1 s}(\mathbf{1}) \mathbf{1} \mathbf{s}(\mathbf{2})$ is not antisymmetric because $\mathbf{1 s ( 1 ) 1 \overline { s } ( 2 ) \neq - 1 s ( 2 ) 1 \overline { s } ( 1 )}$
(note $1 s(1) 1 \bar{s}(2)$ is shorthand for $1 s(1) 1 s(2) \alpha(1) \beta(2))$

But the linear combination

$$
\begin{aligned}
& \Phi=\frac{1}{\sqrt{2}}(1 s(1) 1 \bar{s}(2)-1 s(2) 1 \bar{s}(1)) \text { is antisymmetric because } \\
& \frac{1}{\sqrt{2}}(1 s(2) 1 \bar{s}(1)-1 s(1) 1 \bar{s}(2))=-\frac{1}{\sqrt{2}}(1 s(1) 1 \bar{s}(2)-1 s(2) 1 \bar{s}(1))=-\Phi
\end{aligned}
$$



$$
\Phi=\frac{1}{\sqrt{2}}(1 s(1) 1 \bar{s}(2) 2 s(3)-1 s(2) 1 \bar{s}(1) 2 s(3))
$$

is antisymmetric wrt exchange of electrons 1 and 2, but not 1 and 3 or 2 and 3 .

$$
\Phi=\frac{1}{\sqrt{6}}\binom{1 s(1) 1 \bar{s}(2) 2 s(3)-1 s(1) 1 \bar{s}(3) 2 s(2)-1 s(2) 1 \bar{s}(1) 2 s(3)}{+1 s(2) 1 \bar{s}(3) 2 s(1)+1 s(3) 1 \bar{s}(1) 2 s(2)-1 s(3) 1 \bar{s}(2) 2 s(1)}
$$

is antisymmetric wrt exchange of all three.

The antisymmetric wavefunction can be written as a Slater determinant
$\mathrm{He}: \quad \Phi=\frac{1}{\sqrt{2}}(1 s(1) 1 \bar{s}(2)-1 s(2) 1 \bar{s}(1))=\frac{1}{\sqrt{2}}\left|\begin{array}{ll}1 s(1) & 1 \bar{s}(1) \\ 1 s(2) & 1 \bar{s}(2)\end{array}\right|$

Li: $\quad \Phi=\frac{1}{\sqrt{6}}\left|\begin{array}{lll}1 s(1) & 1 \bar{s}(1) & 2 s(1) \\ 1 s(2) & 1 \bar{s}(2) & 2 s(2) \\ 1 s(3) & 1 \bar{s}(3) & 2 s(3)\end{array}\right|$

He (triplet excited state, $\mathbf{1 s}^{\mathbf{1}} \mathbf{2 s}^{\mathbf{1}}$ ):
$\Phi=\frac{1}{\sqrt{2}}\left|\begin{array}{ll}1 s(1) & 2 s(1) \\ 1 s(2) & 2 s(2)\end{array}\right|=\frac{1}{\sqrt{2}}(1 s(1) 2 s(2)-1 s(2) 2 s(1))$
Note using a single Slater determinant to represent the wavefunction is a convenient way to ensure that it conforms to the Pauli Principle, but it is by no means the only way (any linear combination of Slater determinants will also do the job, see configuration interaction)
(n.b. only 1 electron, so Slater "determinant" is just a $1 \times 1$ determinant: $\Phi=|\boldsymbol{\psi}|$ )

$$
\begin{aligned}
& \psi_{\text {trial }}=\sum_{n} c_{n} \chi_{n}=c_{a} \chi_{1 s a}+c_{b} \chi_{1 s b} \\
& E_{\text {trial }}=\frac{\langle\psi| \hat{\boldsymbol{H}}|\psi\rangle}{\langle\psi \mid \psi\rangle}=\frac{\left\langle\left(c_{a} \chi_{1 s a}+c_{b} \chi_{1 s b}\right)\right| \hat{\boldsymbol{H}}\left|\left(c_{a} \chi_{1 s a}+c_{b} \chi_{1 s b}\right)\right\rangle}{\left\langle\left(c_{a} \chi_{1 s a}+c_{b} \chi_{1 s b}\right) \mid\left(c_{a} \chi_{1 s a}+c_{b} \chi_{1 s b}\right)\right\rangle}
\end{aligned}
$$

Following the process set out in Prof Brouard's 'Valence' lecture 4 (handout pp 18-23) or the equivalent derivation given in Appendix 1, minimisation of the trial wavefunction gives the following secular determinant:

$$
\operatorname{det}|\boldsymbol{H}-\boldsymbol{E S}|=\left|\begin{array}{ll}
\boldsymbol{H}_{a a}-\boldsymbol{E} \boldsymbol{S}_{a a} & \boldsymbol{H}_{a b}-\boldsymbol{E} \boldsymbol{S}_{a b} \\
\boldsymbol{H}_{a b}-\boldsymbol{E} \boldsymbol{S}_{a b} & \boldsymbol{H}_{b b}-\boldsymbol{E} \boldsymbol{S}_{b b}
\end{array}\right|=0
$$

with eigenvalues that we can formulate in terms of matrix elements

$$
\boldsymbol{E}_{ \pm}=\frac{\left(\boldsymbol{H}_{a a} \pm \boldsymbol{H}_{a b}\right)}{\left(\mathbf{1} \pm \boldsymbol{S}_{a b}\right)} \text { and } \boldsymbol{H}_{a a}=\left\langle\boldsymbol{\chi}_{1 s a}\right| \hat{\boldsymbol{H}}\left|\boldsymbol{\chi}_{1 s a}\right\rangle \text { etc. }
$$

(but note we have not yet defined $\hat{\boldsymbol{H}}$ !)
and eigenfunctions that we can anticipate from symmetry considerations:

$$
\psi_{ \pm}=1 \sigma_{g / u}=\frac{1}{\sqrt{2\left(1 \pm S_{a b}\right)}}\left(\chi_{1 s a} \pm \chi_{1 s b}\right)
$$

To make further progress we need to convert the abstract " $\boldsymbol{H}_{a a}$ ", " $\boldsymbol{H}_{a b}$ " etc. into hard numbers.

$$
\begin{gathered}
\widehat{H}=-\frac{\hbar^{2}}{2 m} \nabla^{2}+V=-\frac{\hbar^{2}}{2 m} \nabla^{2}-\frac{e^{2}}{4 \pi \varepsilon_{0}}\left(\frac{1}{r_{a}}\right)-\frac{e^{2}}{4 \pi \varepsilon_{0}}\left(\frac{1}{r_{b}}\right)+\frac{e^{2}}{4 \pi \varepsilon_{0}}\left(\frac{1}{R_{a b}}\right) \\
=-\frac{\hbar^{2}}{2 m} \nabla^{2}-\frac{j_{0}}{r_{a}}-\frac{j_{0}}{r_{b}}+\frac{j_{0}}{R_{a b}} \\
j_{0}=\frac{e^{2}}{4 \pi \varepsilon_{0}}
\end{gathered}
$$


$\frac{j_{0}}{R_{a b}}$ is the repulsion between two nuclei - it is independent of the electron coordinates, so adds a constant term to the total energy. We will neglect this term in the following derivations (it can be trivially added in at the end of any energy calculation).

## A note on units:

The Hamiltonian

$$
\widehat{H}=-\frac{\hbar^{2}}{2 m} \nabla^{2}+V=-\frac{\hbar^{2}}{2 m} \nabla^{2}-\frac{e^{2}}{4 \pi \varepsilon_{0}}\left(\frac{1}{r_{a}}\right)-\frac{e^{2}}{4 \pi \varepsilon_{0}}\left(\frac{1}{r_{b}}\right)+\frac{e^{2}}{4 \pi \varepsilon_{0}}\left(\frac{1}{R_{a b}}\right)
$$

is given in SI units (distance in metres, mass in kg, charge in coulombs, energy in Joules).

It is also often expressed in atomic units (distance in multiples of $\mathrm{a}_{0}$ [the Bohr radius], mass in multiples of $m_{e}$, charge in multiples of e, permittivity in multiples of $4 \pi \varepsilon_{0}$, energy in multiples of Hartrees (atomic units).

$$
\begin{gathered}
\mathrm{a}_{0}=0.529 \times 10^{-10} \mathrm{~m}=0.529 \AA \quad \mathrm{e}=1.6022 \times 10^{-19} \mathrm{C} \quad \mathrm{~m}_{\mathrm{e}}=9.1095 \times 10^{-31} \mathrm{~kg} \\
1 \text { hartree }=1 \mathrm{au}=\frac{\hbar^{2}}{m_{e} a_{0}^{2}}=4.3598 \times 10^{-18} \mathrm{~J}=27.21 \mathrm{eV}
\end{gathered}
$$

In which case the Hamiltonian simplifies to $\quad \widehat{H}=-\frac{1}{2} \nabla^{2}-\frac{1}{r_{a}}-\frac{1}{r_{b}}+\frac{1}{R_{a b}}$

See also 'Valence' lecture 2, P 8:
Coulomb integral: $\quad \alpha=\boldsymbol{H}_{a a}=\left\langle\chi_{a}\right| \hat{\boldsymbol{H}}\left|\chi_{a}\right\rangle=\left\langle\chi_{a}\right|-\frac{\mathbf{1}}{\mathbf{2}} \nabla^{2}-\frac{\mathbf{1}}{r_{a}}-\frac{1}{r_{b}}\left|\chi_{a}\right\rangle$
We will see how to compute these integrals shortly, but for now just note that $\left\langle\chi_{a}\right|-\frac{1}{2} \nabla^{2}-\frac{1}{r_{A}}\left|\chi_{a}\right\rangle=E_{1 s}$, the energy of an electron in a hydrogen atom.

$$
\therefore \alpha=E_{1 s}-\left\langle\chi_{a}\right| \frac{\mathbf{1}}{r_{b}}\left|\chi_{a}\right\rangle=E_{1 s}-j^{\prime}
$$

$j^{\prime}=\left\langle\chi_{a}\right| \frac{\mathbf{1}}{r_{b}}\left|\chi_{a}\right\rangle=$ Coulomb attraction between charge density $\chi_{a}{ }^{2}$ and nucleus b



Likewise, the resonance integral: $\boldsymbol{\beta}=\boldsymbol{H}_{a b}=\left\langle\chi_{a}\right| \hat{\boldsymbol{H}}\left|\chi_{b}\right\rangle=\left\langle\chi_{a}\right|-\frac{\mathbf{1}}{\mathbf{2}} \nabla^{2}-\frac{\mathbf{1}}{\boldsymbol{r}_{a}}-\frac{\mathbf{1}}{\boldsymbol{r}_{b}}\left|\chi_{b}\right\rangle$
but $\left(-\frac{1}{2} \nabla^{2}-\frac{1}{r_{b}}\right)\left|\chi_{b}\right\rangle=\boldsymbol{E}_{1 s}\left|\chi_{b}\right\rangle$
$\therefore \beta=\left\langle\chi_{a} \mid \chi_{b}\right\rangle E_{1 s}-\left\langle\chi_{a}\right| \frac{1}{r_{a}}\left|\chi_{b}\right\rangle=S E_{1 s}-\boldsymbol{k}^{\prime}$
$\boldsymbol{k}^{\prime}=\left\langle\chi_{a}\right| \frac{\mathbf{1}}{r_{a}}\left|\chi_{b}\right\rangle=$ interaction between overlap charge density $\chi_{a} \chi_{b}$ and nucleus a


Sub into solutions of secular equation:

$$
E_{1 \sigma_{g} / 1 \sigma_{u}}=\frac{(\alpha \pm \beta)}{(1 \pm S)}=\frac{\left(E_{1 s}-j^{\prime} \pm\left(S E_{1 s}-k^{\prime}\right)\right)}{(1 \pm S)} \quad E_{1 \sigma_{g}}=E_{1 s}-\frac{\left(j^{\prime}+k^{\prime}\right)}{1+S} \text { or } E_{1 \sigma_{u}}=E_{1 s}-\frac{\left(j^{\prime}-k^{\prime}\right)}{1-S}
$$

## Basis functions

In order to evaluate $\boldsymbol{E}_{1 \sigma_{g}}$ or $\boldsymbol{E}_{1 \sigma_{u}}$, we need to be able to calculate the various integrals involved in the above expressions. These include the overlap integrals, $\langle\chi \mid \chi\rangle$, and the components of $\boldsymbol{H}_{i j}$, which are the kinetic energy, $\langle\chi|-\frac{1}{2} \nabla^{2}|\chi\rangle$, and the electron-nucleus interaction, $\langle\chi| \frac{\mathbf{1}}{\boldsymbol{r}_{\boldsymbol{A}}}|\chi\rangle$. When we get to $\mathrm{H}_{2}$, we will find that we also need to calculate electron-electron repulsion integrals of the type $\langle\chi(\mathbf{1}) \chi(\mathbf{2})| \frac{\mathbf{1}}{r_{12}}|\chi(\mathbf{1}) \chi(\mathbf{2})\rangle$

To calculate these numbers we need to choose a mathematical representation of the atomic orbitals, $\chi$. For discrete molecules, there are 2 plausible choices, Slater-type orbitals and Gaussian-type orbitals. For solid-state (periodic) calculations, there is a third choice, a basis set of plane waves, but we will not be concerned with that here.

## Slater-type orbitals (STOs):

$$
S T O=N_{1} r^{n-1} e^{-\zeta r} Y_{l m}(\theta, \phi)
$$

$N_{1}$ is the normalising constant.
$Y_{l m}(\theta, \phi)$ is a spherical harmonic, controlling angular dependence
$\zeta$ is a measure of how contracted the function is (larger $\zeta$ implies more contracted). In very simple (minimal) basis sets, $\zeta$ is related to Slater's effective nuclear charge, which controls radial behaviour, and can be calculated using Slater's simple empirical rules.

These replicate the behaviour of real atomic orbitals, so are the most intuitive option.

## Gaussian-type orbitals (GTOs): $\quad G T O=N_{1} x^{i} y^{j} z^{k} e^{-\alpha r^{2}}=N_{1} x^{i} y^{j} z^{k} e^{-\alpha\left(x^{2}+y^{2}+z^{2}\right)}$

Angular properties are specified by $i, j, k:(0,0,0)=s ;(1,0,0)=p_{x} ;(1,1,0)=d_{x y}$ etc etc.
The major difference is that the exponential decay is $e^{-\alpha r^{2}}$ and not $e^{-\alpha r}$, which means that the orbital decays too quickly at large $r$, and fails to reproduce the cusp at the nucleus found in real atomic orbitals. Thus they are less accurate representations of a real atomic orbital than Slater functions.


However, GTOs have 2 major advantages:
(1) GTOs are separable in the $x, y$ and $z$ directions:

$$
G T O=N_{1} x^{i} y^{j} z^{k} e^{-\alpha\left(x^{2}+y^{2}+z^{2}\right)}=N_{1}\left(x^{i} e^{-\alpha x^{2}}\right)\left(y^{j} e^{-\alpha y^{2}}\right)\left(z^{k} e^{-\alpha z^{2}}\right)
$$

the same cannot be said of STOs:

$$
S T O=N_{1} e^{-\zeta \sqrt{\left(x^{2}+y^{2}+z^{2}\right)}} \neq N_{1}\left(e^{-\zeta x}\right)\left(e^{-\zeta y}\right)\left(e^{-\zeta z}\right)
$$

(2) the product of 2 GTOs on different centres is just a different GTO centred somewhere in between (the Gaussian Product Theorem).

Consider the product of two s-type gaussians ( $i=j=k=0$ ) with equal exponents, $\alpha$, one centred at $\boldsymbol{x}=\mathbf{0}$, the other at $\boldsymbol{x}=\boldsymbol{a}$

$$
G T O 1=N_{1} e^{-\alpha x^{2}} \quad G T O 2=N_{2} e^{-\alpha(x-a)^{2}}
$$

$$
\begin{aligned}
G T O 1 \times G T O 2 & =N_{1} N_{2} e^{-\alpha x^{2}} e^{-\alpha(x-a)^{2}}=N_{1} N_{2} e^{-\alpha\left[x^{2}+(x-a)^{2}\right]}=N_{1} N_{2} e^{-\alpha\left(2 x^{2}-2 a x+a^{2}\right)} \\
& =N_{1} N_{2} e^{-2 \alpha\left(x^{2}-a x+\frac{a^{2}}{2}\right)}=N_{1} N_{2} e^{-2 \alpha\left[\left(x-\frac{a}{2}\right)^{2}+\frac{a^{2}}{4}\right]}=e^{-\frac{\alpha a^{2}}{2}} N_{1} N_{2} e^{-2 \alpha\left(x-\frac{a}{2}\right)^{2}}
\end{aligned}
$$

i.e. the product is another Gaussian with exponent $2 \alpha$ centred half way between the original basis functions ( $\boldsymbol{x}=\frac{\boldsymbol{a}}{\mathbf{2}}$ ) and scaled by $e^{-\frac{\alpha a^{2}}{2}}$


Figure 3.1 The product of two $1 s$ Gaussians is a third $1 s$ Gaussian.

These two features combine to make the computation of all integrals much easier for Gaussian functions than for Slater functions. It is harder, but still possible, to calculate the 1electron integrals (overlap, kinetic energy, electron-nucleus interaction) using a Slater-type basis, but if electron-electron repulsion integrals are needed (as they are for anything with >1 electron!), Slater-type functions are a very poor choice.

Example 1: an overlap integral between two s orbitals on different atoms.

Take 2 s-type gaussians with equal exponents, one at the origin, the other at $(0,0, a)$.

$$
S=N_{1} N_{2} \iiint_{-\infty}^{+\infty} e^{-\alpha\left(x^{2}+y^{2}+z^{2}\right)} e^{-\alpha\left(x^{2}+y^{2}+(z-a)^{2}\right)} d x d y d z
$$

First, we need to normalise each basis function.

$$
\begin{aligned}
1=N_{1}^{2} \iiint_{-\infty}^{+\infty} & e^{-\alpha\left(x^{2}+y^{2}+z^{2}\right)} e^{-\alpha\left(x^{2}+y^{2}+z^{2}\right)} d x d y d z \\
& =N_{1}^{2} \int_{-\infty}^{+\infty} e^{-2 \alpha x^{2}} d x \int_{-\infty}^{+\infty} e^{-2 \alpha y^{2}} d y \int_{-\infty}^{+\infty} e^{-2 \alpha z^{2}} d z
\end{aligned}
$$

using the standard integral

$$
\begin{gathered}
\int_{-\infty}^{+\infty} e^{-b x^{2}} d x=\sqrt{\frac{\pi}{b}} \\
1=N_{1}^{2}\left(\frac{\pi}{2 \alpha}\right)^{1 / 2}\left(\frac{\pi}{2 \alpha}\right)^{1 / 2}\left(\frac{\pi}{2 \alpha}\right)^{1 / 2}=N_{1}^{2}\left(\frac{\pi}{2 \alpha}\right)^{3 / 2}
\end{gathered}
$$

$$
\begin{gathered}
N_{1}\left(=N_{2}\right)=\left(\frac{2 \alpha}{\pi}\right)^{3 / 4} \\
S=\left(\frac{2 \alpha}{\pi}\right)^{3 / 4} \cdot\left(\frac{2 \alpha}{\pi}\right)^{3 / 4} \int_{-\infty}^{+\infty} e^{-2 \alpha x^{2}} d x \int_{-\infty}^{+\infty} e^{-2 \alpha y^{2}} d y \int_{-\infty}^{+\infty} e^{-\alpha\left[z^{2}+(z-a)^{2}\right]} d z \\
=\left(\frac{2 \alpha}{\pi}\right)^{3 / 2}\left(\frac{\pi}{2 \alpha}\right)^{1 / 2}\left(\frac{\pi}{2 \alpha}\right)^{1 / 2} \int_{-\infty}^{+\infty} e^{-\alpha\left[z^{2}+(z-a)^{2}\right]} d z
\end{gathered}
$$

Using the Gaussian product theorem to deal with the integral in $z$ :

$$
\begin{gathered}
\int_{-\infty}^{+\infty} e^{-\alpha\left[z^{2}+(z-a)^{2}\right]} d z=e^{-\alpha\left(\frac{a^{2}}{2}\right)} \int_{-\infty}^{+\infty} e^{-2 \alpha\left(z-\frac{a}{2}\right)^{2}} d z=e^{-\alpha\left(\frac{a^{2}}{2}\right)}\left(\frac{\pi}{2 \alpha}\right)^{1 / 2} \\
\therefore S=\left(\frac{2 \alpha}{\pi}\right)^{3 / 2}\left(\frac{\pi}{2 \alpha}\right)^{1 / 2}\left(\frac{\pi}{2 \alpha}\right)^{1 / 2}\left(\frac{\pi}{2 \alpha}\right)^{1 / 2} e^{-\alpha\left(\frac{a^{2}}{2}\right)}=e^{-\alpha\left(\frac{a^{2}}{2}\right)}
\end{gathered}
$$

Example 2: kinetic energy of an electron in a 1s orbital, $\left\langle\phi_{1 s}\right| \hat{T}\left|\phi_{1 s}\right\rangle$

$$
\widehat{T}=-\frac{1}{2} \nabla^{2}=-\frac{1}{2}\left(\frac{\delta^{2}}{\delta x^{2}}+\frac{\delta^{2}}{\delta y^{2}}+\frac{\delta^{2}}{\delta y^{2}}\right)
$$

Taking the $x$ component ( $y$ and $z$ give identical contributions), we need to evaluate:

$$
\begin{gathered}
N_{1} N_{2} \int_{-\infty}^{+\infty} e^{-\alpha x^{2}} \frac{d^{2}}{d x^{2}}\left(e^{-\alpha x^{2}}\right) d x \int_{-\infty}^{+\infty} e^{-2 \alpha y^{2}} d y \int_{-\infty}^{+\infty} e^{-2 \alpha z^{2}} d z \\
\frac{d^{2}}{d x^{2}}\left(e^{-\alpha x^{2}}\right)=\left(4 \alpha^{2} x^{2}-2 \alpha\right) e^{-\alpha x^{2}}
\end{gathered}
$$

So we need to evaluate 2 integrals and sum them:

$$
-2 \alpha N_{1} N_{2} \int_{-\infty}^{+\infty} e^{-2 \alpha x^{2}} d x \int_{-\infty}^{+\infty} e^{-2 \alpha y^{2}} d y \int_{-\infty}^{+\infty} e^{-2 \alpha z^{2}} d z
$$

and

$$
4 \alpha^{2} N_{1} N_{2} \int_{-\infty}^{+\infty} x^{2} e^{-2 \alpha x^{2}} d x \int_{-\infty}^{+\infty} e^{-2 \alpha y^{2}} d y \int_{-\infty}^{+\infty} e^{-2 \alpha z^{2}} d z
$$

(for which we need the standard integral $\int_{-\infty}^{+\infty} x^{2} e^{-b x^{2}} d x=\frac{1}{2 b} \sqrt{\frac{\pi}{b}}$ )

Attraction (to nuclei) and repulsion (with other electrons) integrals are trickier because of the $1 / r$ terms, for which we need Fourier transform techniques. However, relatively simple closed analytical forms exist. The same is not true for Slater-type functions!

## Practical choices of basis set.

Despite their mathematical convenience, it remains true that GTOs provide a much worse approximation to atomic orbitals than STOs. The compromise is to use a linear combination of several GTOs to represent each atomic orbital rather than a single STO. For example, a triple-zeta basis set uses three GTOs to describe each atomic orbital. The basic philosophy is that it is easier to perform a large number of easy integrals than to perform a much small number of hard ones!


If we vary the coefficients in the sum, we can get an orbital of arbitrary size.


And we can introduce radial nodes by using negative coefficients (the choices below are arbitrary, simply designed to illustrate how you can create different radial distribution functions from the same set of gaussian functions.


For example an STO-3G basis set uses 3 GTOs to represent each STO. Thus a calculation on $\mathrm{H}_{2} \mathrm{O}$ with an STO-3G basis would involve 21 basis functions ( 3 for each of $\mathrm{O} 1 \mathrm{~s}, 2 \mathrm{~s}, 2 \mathrm{p}_{\mathrm{x}}$ and $\mathrm{H} 1 \mathrm{~s}(\mathrm{x} 2)$ ).
Very large basis sets are commonly used now, including:

1) Double, triple, quadruple.... zeta forms: more than one exponent is used to describe a given orbital. Varying linear combinations allow radial extent of orbital to vary. e.g 6-31G
2) Polarisation functions: basis functions with higher angular momentum (e.g. psymmetry functions on H , d-symmetry functions on B,C,N,O,F, f-symmetry functions on transition metals) e.g 6-31G* (recall the discussion of the role of $d$ orbitals in e.g. $\mathrm{SF}_{6}$ )
3) Diffuse functions: very low zeta - important for accurate description of weakly bound electrons (anions, for example) e.g 6-31+G
Choice of basis is typically one of the major decisions practising computational chemists need to make.
$\mathrm{H}_{2}$ : explicit formulation of the Hamiltonian for a 2-electron system
$\hat{\boldsymbol{H}}=\sum_{i}\left(-\frac{1}{2} \nabla_{i}{ }^{2}+V_{i}\right)=\left(-\frac{1}{2} \nabla_{1}{ }^{2}-\frac{1}{r_{a 1}}-\frac{1}{r_{b 1}}\right)+\left(-\frac{1}{2} \nabla_{2}{ }^{2}-\frac{1}{r_{a 2}}-\frac{1}{r_{b 2}}\right)+\frac{1}{r_{12}}=\hat{\boldsymbol{H}}_{1}+\hat{\boldsymbol{H}}_{2}+\frac{1}{r_{12}}$

Note $\hat{\boldsymbol{H}}_{1}$ and $\hat{\boldsymbol{H}}_{2}$ are identical to the 1-electron Hamiltonians for $\mathrm{H}_{2}{ }^{+}$
if $\frac{1}{r_{12}}$ was ignored completely (i.e. assuming electrons don't interact), this is simply the sum of two independent one-electron $\mathrm{H}_{2}^{+}$Hamiltonians, and the problem is separable. The resulting energy and ground-state wavefunction would be

$$
E=\mathbf{2} E_{1 \sigma_{g}} \quad \Psi=\mathbf{1} \sigma_{g}(\mathbf{1}) \mathbf{1} \bar{\sigma}_{g}(\mathbf{2}) \quad \text { (a simple "Hartree product") }
$$

But we have established that a wavefunction of this type is not antisymmetric wrt exchange of electrons (unsurprisingly as we deliberately ignored the interaction between the electrons in deriving it!): we need to use a Slater determinant
$\Phi=\left|\mathbf{1} \sigma_{g} 1 \bar{\sigma}_{g}\right|=\frac{1}{\sqrt{2}}\left|\begin{array}{ll}\mathbf{1} \sigma_{g}(\mathbf{1}) & \mathbf{1} \bar{\sigma}_{g}(\mathbf{1}) \\ 1 \sigma_{g}(\mathbf{2}) & \mathbf{1} \bar{\sigma}_{g}(\mathbf{2})\end{array}\right|=\frac{1}{\sqrt{2}}\left(\mathbf{1} \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})-\mathbf{1} \sigma_{g}(\mathbf{2}) \mathbf{1} \bar{\sigma}_{g}(\mathbf{1})\right)$
Now we have an expression for $\Phi$ and an expression for $\hat{\boldsymbol{H}}$, we can compute the expectation value:

$$
\langle\Phi| \hat{\boldsymbol{H}}|\Phi\rangle=\boldsymbol{E}
$$

Step 1: $\quad$ Expand $\Phi$ as a Slater determinant

$$
\begin{aligned}
&\langle\Phi| \hat{H}|\Phi\rangle=E=\frac{1}{2}\left\langle\left(1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(\mathbf{2})-\mathbf{1} \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right)\right| \hat{H}_{1}+\hat{H}_{2}+\frac{1}{r_{12}}\left|\left(1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(\mathbf{2})-\mathbf{1} \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right)\right\rangle \\
& E=\frac{1}{2}\left\langle 1 \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right| \hat{H}_{1}+\hat{H}_{2}+\frac{1}{r_{12}}\left|1 \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right\rangle \\
&-\frac{1}{2}\left\langle 1 \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right| \hat{H}_{1}+\hat{H}_{2}+\frac{1}{r_{12}}\left|1 \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right\rangle \\
&-\frac{1}{2}\left\langle 1 \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right| \hat{H}_{1}+\hat{H}_{2}+\frac{1}{r_{12}}\left|1 \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right\rangle \\
&+\frac{1}{2}\left\langle 1 \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(1)\right| \hat{H}_{1}+\hat{H}_{2}+\frac{1}{r_{12}}\left|1 \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right\rangle
\end{aligned}
$$

Step 2: Expand terms of $\hat{\boldsymbol{H}}$ to give 12 integrals, $\boldsymbol{I}_{1-12}$

$$
\begin{aligned}
& E=\frac{1}{2}\left\langle 1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right| \hat{H}_{1}\left|1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right\rangle+\frac{1}{2}\left\langle 1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right| \hat{H}_{2}\left|1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right\rangle \\
& -\frac{1}{2}\left\langle 1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right| \hat{H}_{1}\left|1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right\rangle-\frac{1}{2}\left\langle 1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right| \hat{H}_{2}\left|1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right\rangle \\
& -\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right| \hat{H}_{1}\left|1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right\rangle-\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right| \hat{H}_{2}\left|1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right\rangle \\
& +\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right| \hat{H}_{1}\left|1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right\rangle+\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right| \hat{H}_{2}\left|1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right\rangle \\
& +\frac{1}{2}\left\langle\mathbf{1} \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\mathbf{1} \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right\rangle-\frac{1}{2}\left\langle 1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(\mathbf{2})\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right\rangle \\
& -\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right\rangle+\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right\rangle
\end{aligned}
$$

Now look at each of $I_{1-12}$ in turn:

1-electron integrals, $I_{1-8}$
$\boldsymbol{I}_{1}=\left\langle\mathbf{1} \sigma_{g}(\mathbf{1}) \mathbf{1} \bar{\sigma}_{g}(\mathbf{2})\right| \hat{H}_{1}\left|\mathbf{1} \sigma_{g}(\mathbf{1}) \mathbf{1} \bar{\sigma}_{g}(\mathbf{2})\right\rangle=\left\langle\mathbf{1} \sigma_{g}(\mathbf{1})\right| \hat{H}_{1}\left|\mathbf{1} \sigma_{g}(\mathbf{1})\right\rangle\left\langle\mathbf{1} \bar{\sigma}_{g}(\mathbf{2}) \mid \mathbf{1} \bar{\sigma}_{g}(\mathbf{2})\right\rangle$ (note $\hat{\boldsymbol{H}}_{1}$ only acts on electron 1)

Separating the spatial and spin components of the wavefunctions gives

$$
\begin{aligned}
& \boldsymbol{I}_{1}=\left\langle\mathbf{1} \sigma_{g}(\mathbf{1})\right| \hat{H}_{1}\left|\mathbf{1} \sigma_{g}(\mathbf{1})\right\rangle\langle\alpha(\mathbf{1}) \mid \alpha(\mathbf{1})\rangle\left\langle\mathbf{1} \sigma_{g} \mathbf{( 2 )} \mid \mathbf{1} \sigma_{g}(\mathbf{2})\right\rangle\langle\beta(\mathbf{2}) \mid \beta(\mathbf{2})\rangle \\
& =\quad \boldsymbol{E}_{1 \sigma_{g}} \quad \times 1 \times 1 \times 1 \quad=\boldsymbol{E}_{1 \sigma_{g}}
\end{aligned}
$$

Similarly

$$
\begin{aligned}
& I_{2}=\left\langle 1 \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right| \hat{H}_{2}\left|1 \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right\rangle=E_{1 \sigma_{g}} \\
& I_{7}=\left\langle 1 \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right| \hat{H}_{1}\left|1 \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(1)\right\rangle=E_{1 \sigma_{g}} \\
& I_{8}=\left\langle 1 \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right| \hat{H}_{2}\left|1 \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right\rangle=E_{1 \sigma_{g}}
\end{aligned}
$$


(due to spin orthogonality)

Similarly

$$
\begin{aligned}
& I_{4}=\left\langle 1 \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right| \hat{H}_{2}\left|\mathbf{1} \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right\rangle=0 \\
& I_{5}=\left\langle\mathbf{1} \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right| \hat{H}_{1}\left|1 \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right\rangle=0 \\
& I_{6}=\left\langle 1 \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right| \hat{H}_{2}\left|1 \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right\rangle=0
\end{aligned}
$$

2-electron integrals (terms in $1 / \mathrm{r}_{12}$ ), $I_{9-12}$
$\boldsymbol{I}_{9}=\left\langle\mathbf{1} \sigma_{g}(\mathbf{1}) \mathbf{1} \bar{\sigma}_{g}(\mathbf{2})\right| \frac{\mathbf{1}}{r_{12}}\left|\mathbf{1} \sigma_{g}(\mathbf{1}) \mathbf{1} \bar{\sigma}_{g}(\mathbf{2})\right\rangle$
Separating spatial and spin components:
$=\left\langle 1 \sigma_{g}(1) 1 \sigma_{g}(2)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(1) 1 \sigma_{g}(2)\right\rangle\langle\alpha(1) \mid \alpha(1)\rangle\langle\beta(2) \mid \beta(2)\rangle$
$=\left\langle 1 \sigma_{g}(1) 1 \sigma_{g}(2)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(1) 1 \sigma_{g}(2)\right\rangle \times 1 \quad \times 1 \quad=J_{1 \sigma_{g} 1 \sigma_{g}}$
This is the 'Coulomb integral' (note different use of the term 'Coulomb integral').
Physically: the repulsion between two electrons in the same orbital, $\mathbf{1} \sigma_{g}$

## Physicists' and chemists' notation

A shorthand notation for the 2-electron integral $\left\langle\mathbf{1} \sigma_{g} \mathbf{( 1 ) 1} \bar{\sigma}_{g}(\mathbf{2})\right| \frac{\mathbf{1}}{\boldsymbol{r}_{\mathbf{1 2}}}\left|\mathbf{1} \boldsymbol{\sigma}_{g}(\mathbf{1}) \mathbf{1} \bar{\sigma}_{g}(\mathbf{2})\right\rangle$ is $\left\langle\mathbf{1} \sigma_{g}(\mathbf{1}) \mathbf{1} \bar{\sigma}_{g}(\mathbf{2}) \mid \mathbf{1} \sigma_{g}(\mathbf{1}) \mathbf{1} \bar{\sigma}_{g}(\mathbf{2})\right\rangle$ or more generally $\langle\boldsymbol{i} \mid \boldsymbol{i}\rangle$ where the $1 / r_{12}$ term is implicit. This is called the physicists notation. Up to 4 spin orbitals can be involved, in which case the integral is denoted $\langle\boldsymbol{i} \boldsymbol{j} \mid \boldsymbol{k} \boldsymbol{l}\rangle$.
It is often conceptually easier to collect all terms relating to a given electron on the same side of the $1 / r_{12}$ term.
$\left\langle 1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right\rangle=\int 1 \sigma_{g}(1) * 1 \bar{\sigma}_{g}(2) * \frac{1}{r_{12}} 1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)=$
$\int 1 \sigma_{g}(1) * 1 \sigma_{g}(1) \frac{1}{r_{12}} 1 \bar{\sigma}_{g}(2) * 1 \bar{\sigma}_{g}(2)=[i i \mid j j]$ This is called chemists' notation.
Note the use of a square bracket rather than a traditional 'bra'/'ket'. This is because the complex conjugates have moved (complex conjugation is implicit in ' $\langle |$ ').

## Returning to the expansion:

$I_{12}=\left\langle 1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right\rangle=J_{1 \sigma_{g} 1 \sigma_{g}}=I_{9}$
but $\quad \boldsymbol{I}_{10}=\left\langle\mathbf{1} \sigma_{g}(\mathbf{1}) \mathbf{1} \bar{\sigma}_{g}(\mathbf{2})\right| \frac{\mathbf{1}}{r_{12}}\left|\mathbf{1} \sigma_{g}(\mathbf{2}) \mathbf{1} \bar{\sigma}_{g}(\mathbf{1})\right\rangle$
$=\left\langle 1 \sigma_{g}(1) 1 \sigma_{g}(2)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(1) 1 \sigma_{g}(2)\right\rangle\langle\alpha(1) \mid \beta(1)\rangle\langle\beta(2) \mid \alpha(2)\rangle$
$=\left\langle 1 \sigma_{g}(\mathbf{1}) 1 \sigma_{g}(\mathbf{2})\right| \frac{\mathbf{1}}{r_{12}}\left|\mathbf{1} \sigma_{g}(\mathbf{1}) 1 \sigma_{g}(\mathbf{2})\right\rangle \quad \mathrm{x} 0 \quad \mathrm{x} 0 \quad=0$
Similarly $\boldsymbol{I}_{11}=\left\langle 1 \sigma_{g}(\mathbf{2}) 1 \bar{\sigma}_{g}(\mathbf{1})\right| \frac{1}{r_{12}}\left|\mathbf{1} \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right\rangle=\mathbf{0}$
So $E=\frac{1}{2}\left(4 E_{1 \sigma_{g}}+2 J_{1 \sigma_{g} 1 \sigma_{g}}\right)=2 E_{1 \sigma_{g}}+J_{1 \sigma_{g} 1 \sigma_{g}}$
i.e. twice the energy of $\mathrm{H}_{2}{ }^{+}$, plus an addition term $J_{1 \sigma_{g} 1 \sigma_{g}}$ for the electron-electron repulsion.

## Excited states of $\mathrm{H}_{2}$ and the exchange integral, K.

Let us perform the same analysis with the first triplet excited state of $\mathrm{H}_{2}$,
$\Phi=\left|1 \sigma_{g} 1 \sigma_{u}\right|=\frac{1}{\sqrt{2}}\left|\begin{array}{ll}1 \sigma_{g}(1) & 1 \sigma_{u}(1) \\ 1 \sigma_{g}(2) & 1 \sigma_{u}(2)\end{array}\right|=\frac{1}{\sqrt{2}}\left(1 \sigma_{g}(1) 1 \sigma_{u}(2)-1 \sigma_{g}(2) 1 \sigma_{u}(1)\right)$
Expanding $\Phi$ and $\hat{\boldsymbol{H}}$ again gives 12 integrals:

$$
\begin{aligned}
& E=\frac{1}{2}\left\langle 1 \sigma_{g}(1) 1 \sigma_{u}(2)\right| \hat{H}_{1}\left|1 \sigma_{g}(1) 1 \sigma_{u}(2)\right\rangle+\frac{1}{2}\left\langle 1 \sigma_{g}(1) 1 \sigma_{u}(2)\right| \hat{H}_{2}\left|1 \sigma_{g}(1) 1 \sigma_{u}(2)\right\rangle \\
& -\frac{1}{2}\left\langle 1 \sigma_{g}(1) 1 \sigma_{u}(2)\right| \hat{H}_{1}\left|1 \sigma_{g}(\mathbf{2}) 1 \sigma_{u}(1)\right\rangle-\frac{1}{2}\left\langle 1 \sigma_{g}(1) 1 \sigma_{u}(2)\right| \hat{H}_{2}\left|1 \sigma_{g}(2) 1 \sigma_{u}(1)\right\rangle \\
& -\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \sigma_{u}(1)\right| \hat{H}_{1}\left|1 \sigma_{g}(1) 1 \sigma_{u}(2)\right\rangle-\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \sigma_{u}(1)\right| \hat{H}_{2}\left|1 \sigma_{g}(1) 1 \sigma_{u}(2)\right\rangle \\
& +\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \sigma_{u}(1)\right| \hat{H}_{1}\left|1 \sigma_{g}(2) 1 \sigma_{u}(1)\right\rangle+\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \sigma_{u}(1)\right| \hat{H}_{2}\left|1 \sigma_{g}(2) 1 \sigma_{u}(1)\right\rangle \\
& +\frac{1}{2}\left\langle 1 \sigma_{g}(1) 1 \sigma_{u}(2)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(1) 1 \sigma_{u}(2)\right\rangle-\frac{1}{2}\left\langle 1 \sigma_{g}(1) 1 \sigma_{u}(2)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(2) 1 \sigma_{u}(1)\right\rangle \\
& -\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \sigma_{u}(1)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(1) 1 \sigma_{u}(2)\right\rangle+\frac{1}{2}\left\langle 1 \sigma_{g}(2) 1 \sigma_{u}(1)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(2) 1 \sigma_{u}(1)\right\rangle
\end{aligned}
$$

$I_{1}=\left\langle 1 \sigma_{g}(1) 1 \sigma_{u}(2)\right| \hat{H}_{1}\left|1 \sigma_{g}(1) 1 \sigma_{u}(2)\right\rangle=I_{8}=\left\langle 1 \sigma_{g}(2) 1 \sigma_{u}(1)\right| \hat{H}_{2}\left|1 \sigma_{g}(2) 1 \sigma_{u}(1)\right\rangle=E_{1 \sigma_{g}}$ $I_{2}=\left\langle 1 \sigma_{g}(1) 1 \sigma_{u}(2)\right| \hat{H}_{2}\left|1 \sigma_{g}(1) 1 \sigma_{u}(2)\right\rangle=I_{7}=\left\langle 1 \sigma_{g}(2) 1 \sigma_{u}(1)\right| \hat{H}_{1}\left|1 \sigma_{g}(2) 1 \sigma_{u}(1)\right\rangle=E_{1 \sigma_{u}}$
$I_{3,4,5,6}=0$
$\boldsymbol{I}_{9}=\left\langle\mathbf{1} \sigma_{g}(\mathbf{1}) 1 \sigma_{u}(\mathbf{2})\right| \frac{\mathbf{1}}{r_{12}}\left|\mathbf{1} \sigma_{g}(\mathbf{1}) 1 \sigma_{u}(\mathbf{2})\right\rangle=\boldsymbol{I}_{12}=\left\langle 1 \sigma_{g}(\mathbf{2}) 1 \sigma_{u}(\mathbf{1})\right| \frac{\dot{j}_{0}}{r_{12}}\left|\mathbf{1} \sigma_{g}(\mathbf{2}) 1 \sigma_{u}(\mathbf{1})\right\rangle=J_{1 \sigma_{g} 1 \sigma_{u}}$
Physically: $\boldsymbol{J}_{1 \sigma_{g} 1 \sigma_{u}}$ is the repulsion between an electron in $\mathbf{1} \sigma_{g}$ and an electron in $\mathbf{1} \sigma_{u}$
thus far, all is exactly as before for the ground state, BUT
$\Lambda_{10}=\left\langle 1 \sigma_{g}(1) 1 \sigma_{u}(2)\right| \frac{\dot{j}_{0}}{r_{12}}\left|1 \sigma_{g}(2) 1 \sigma_{u}(1)\right\rangle=I_{11}=\left\langle 1 \sigma_{g}(2) 1 \sigma_{u}(1)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(1) 1 \sigma_{u}(2)\right\rangle$
$=\left\langle 1 \sigma_{g}(\mathbf{1}) 1 \sigma_{u}(\mathbf{2})\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(2) 1 \sigma_{u}(1)\right\rangle\langle\alpha(1) \mid \alpha(1)\rangle\langle\alpha(2) \mid \alpha(2)\rangle$
$=\left\langle 1 \sigma_{g}(1) 1 \sigma_{u}(2)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(2) 1 \sigma_{u}(1)\right\rangle \quad \times 1 \quad \times 1 \quad \neq 0 \quad$ (no spin orthogonality)
$\left\langle 1 \sigma_{g}(\mathbf{1}) 1 \sigma_{u}(\mathbf{2})\right| \frac{\mathbf{1}}{r_{12}}\left|\mathbf{1} \sigma_{g}(\mathbf{2}) 1 \sigma_{u}(1)\right\rangle=K_{1 \sigma_{g} 1 \sigma_{u}}$
So the $\boldsymbol{I}_{10}$ (and $\boldsymbol{I}_{11}$ ) integrals survive: they are known as exchange integrals, $\boldsymbol{K}$ - a purely quantum phenomenon.

Collecting terms: $\quad E=E_{1 \sigma_{g}}+E_{1 \sigma_{u}}+J_{1 \sigma_{g} 1 \sigma_{u}}-K_{1 \sigma_{g} 1 \sigma_{u}}$
$J$ and $K$ are both positive (to be justified later), so $K$ can be thought of as offsetting some of the electron-electron repulsion captured in $J$ : electrons with parallel spin repel each less than electrons with opposite spins.

Note $\hat{H}$ does not operate on spin degrees of freedom. Thus influence of spin on energy is exerted indirectly via the spatial constraints imposed by the Pauli principle

$$
{ }^{3} \Phi=\frac{1}{\sqrt{2}}\left(1 \sigma_{g}(1) 1 \sigma_{u}(2)-1 \sigma_{g}(2) 1 \sigma_{u}(1)\right)=0 \quad \underset{0}{\text { when }} r_{1}
$$

What are the signs and magnitudes of $J$ and $K$ ? Consider the orbitals:

$1 \sigma_{g}$

$1 \sigma_{u}$
(red is positive, blue is negative)

Now consider the products of orbitals involved in the expressions for $J$ and $K$.



$$
1 \sigma_{g}(1) 1 \sigma_{g}(1)
$$

> 0 everywhere

$1 \sigma_{u}(2) 1 \sigma_{u}(2)$
> 0 everywhere
$1 / r_{12}>0$ everywhere
so: $\quad J$ is the sum of terms that are all positive, so $J$ MUST be positive (no surprise - it is a repulsion, after all!).

## What about K?

$$
\left\langle 1 \sigma_{g}(1) 1 \sigma_{u}(2)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(2) 1 \sigma_{u}(1)\right\rangle=K
$$


$1 \sigma_{g}(1) 1 \sigma_{u}(1)$

$1 \sigma_{u}(2) 1 \sigma_{g}(2)$

The products can be either positive or negative, depending on position in space.
If both electrons on same side of the node: $\quad \mathbf{1} \sigma_{g}(\mathbf{1}) \mathbf{1} \sigma_{u}(\mathbf{2}) \frac{\mathbf{1}}{r_{12}} \mathbf{1} \sigma_{g}(\mathbf{2}) \mathbf{1} \sigma_{u}(\mathbf{1})>\mathbf{0}$
If electrons on opposite sides of the node: $\quad \mathbf{1} \sigma_{g}(\mathbf{1}) \mathbf{1} \sigma_{u}(\mathbf{2}) \frac{\mathbf{1}}{r_{12}} \mathbf{1} \sigma_{g}(\mathbf{2}) \mathbf{1} \sigma_{u}(\mathbf{1})<\mathbf{0}$
$K$ is the summation of some terms that are positive and some terms that are negative, so we cannot, a priori, predict its sign in the same way as we can for $J$.

However, if the electrons are on the same side of the node, $\mathrm{r}_{12}$ is necessarily small, so $\mathbf{1} \sigma_{g}(\mathbf{1}) 1 \sigma_{u}(\mathbf{2}) \frac{1}{r_{12}} 1 \sigma_{g}(2) 1 \sigma_{u}(1)$ is large (and positive as established above).
Whereas if the electrons are on opposite sides of the node, $r_{12}$ is necessarily large( $r$ ), so $\mathbf{1} \sigma_{g}(\mathbf{1}) \mathbf{1} \sigma_{u}(\mathbf{2}) \frac{\mathbf{1}}{r_{12}} \mathbf{1} \sigma_{g}(\mathbf{2}) \mathbf{1} \sigma_{u}(\mathbf{1})$ is small (and negative as established above).

Hence $K$ is positive, but generally smaller than $J$ (in practice $\sim 25 \%$ smaller in cases where there are no radial such as first-row atoms ( $\mathrm{C}, \mathrm{N}, \mathrm{O}, \mathrm{F}$ ), first-row TMs and lanthanides)

## Conclusion: $\boldsymbol{J} \boldsymbol{>} \boldsymbol{K} \boldsymbol{>} \mathbf{0}$ (hence parallel spins favoured)

and note that we can trace this fact all the way back to the fact that the sign in the expansion of the wavefunction is negative:

$$
\Phi=\frac{1}{\sqrt{2}}\left(1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)-1 \sigma_{g}(2) 1 \bar{\sigma}_{g}(1)\right)
$$

If it were positive, parallel-spin electrons would be disfavoured!

## Generalised energy expressions for larger molecules.

We can generalise the expression for the energy to $E=\sum_{i} \mathbf{2} \boldsymbol{E}_{i}+\sum_{i, j}\left(\mathbf{2} \boldsymbol{J}_{i j}-\boldsymbol{K}_{i j}\right)$ for any closedshell molecule, where the summation runs over doubly occupied orbitals

Comment: expanding the ground state of $\mathrm{H}_{2}$ using this expression gives $\boldsymbol{E}=\mathbf{2} \boldsymbol{E}_{1 \sigma_{g}}+\mathbf{2} \boldsymbol{J}_{1 \sigma_{g} 1 \sigma_{g}}-\boldsymbol{K}_{1 \sigma_{g} 1 \sigma_{g}}$ which looks different from expression above, but note $K_{1 \sigma_{g} 1 \sigma_{g}}=\left\langle\mathbf{1} \sigma_{g}(\mathbf{1}) 1 \bar{\sigma}_{g}(\mathbf{2})\right| \frac{\mathbf{1}}{r_{12}}\left|\mathbf{1} \bar{\sigma}_{g}(\mathbf{2}) 1 \sigma_{g}(1)\right\rangle=J_{1 \sigma_{g} 1 \sigma_{g}}$
or in general $\boldsymbol{K}_{i i}=\boldsymbol{J}_{i i}$
so $\boldsymbol{E}=\mathbf{2} E_{1 \sigma_{g}}+\mathbf{2} \boldsymbol{J}_{1 \sigma_{g} 1 \sigma_{g}}-\boldsymbol{K}_{1 \sigma_{g} 1 \sigma_{g}}=\mathbf{2} E_{1 \sigma_{g}}+\boldsymbol{J}_{1 \sigma_{g} 1 \sigma_{g}}$ as on page 12

A simple 'algorithm' that works for any atom/molecule (open or closed shell):

The total energy = sum of the one-electron energies +1 J per pairwise interaction between two electrons $-1 K$ per pairwise interaction between two electrons with parallel spins.

For $\mathrm{He}_{2}$ :

$$
\boldsymbol{E}=2 E_{1 \sigma_{g}}+2 E_{1 \sigma_{u}}+J_{1 \sigma_{g} 1 \sigma_{g}}+4 J_{1 \sigma_{g} 1 \sigma_{u}}+J_{1 \sigma_{u} 1 \sigma_{u}}-\mathbf{2 K} K_{1 \sigma_{g} 1 \sigma_{u}}
$$

## How important is $K$ ?

We have established that $K$ offsets some of the e-e repulsion in $J$, how much is the 'offset' worth? Let's look at 2 limiting cases, both with 4 electrons: $\mathrm{He}_{2}$ and Be

1) $\mathrm{He}_{2}$ : when the electrons ( $\sigma_{\mathrm{g}}$ and $\sigma_{\mathrm{u}}$ ) occupy similar regions of space

$K$ is very large - it reduces the Coulomb repulsion by a factor of $2 / 3$
2) Be: effectively the same problem (4 electrons in 2 orbitals), but now the electrons (1s and 2 s) occupy very different regions of space

$K$ is very small - it only offsets $\sim 5 \%$ of the Coulomb repulsion.
Useful rule of thumb for transition metals: $K \cong 0.25 \times J$ (electrons are in similar regions of space).

## Expansion of $J$ and $K$ in terms of atomic basis:

We can expand $1 \sigma_{g}=\frac{1}{\sqrt{2(1+S)}}\left(\chi_{1 s a}+\chi_{1 s b}\right) \quad 1 \sigma_{u}=\frac{1}{\sqrt{2(1-S)}}\left(\chi_{1 s a}-\chi_{1 s b}\right)$
$J_{1 \sigma_{g} 1 \sigma_{g}}=\left\langle 1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right| \frac{1}{r_{12}}\left|1 \sigma_{g}(1) 1 \bar{\sigma}_{g}(2)\right\rangle=$
$\frac{1}{4(1+S)^{2}}\left\langle\left(\chi_{a}+\chi_{b}\right)(\mathbf{1})\left(\chi_{a}+\chi_{b}\right)(2)\right| \frac{1}{r_{12}}\left|\left(\chi_{a}+\chi_{b}\right)(2)\left(\chi_{a}+\chi_{b}\right)(\mathbf{1})\right\rangle$
$=\frac{1}{4(1+S)^{2}}\left[\left\langle\chi_{a}(\mathbf{1}) \chi_{a}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{a}(\mathbf{2}) \chi_{a}(\mathbf{1})\right\rangle+\left\langle\chi_{a}(\mathbf{1}) \chi_{a}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{a}(\mathbf{2}) \chi_{b}(\mathbf{1})\right\rangle\right.$
$+\left\langle\chi_{a}(\mathbf{1}) \chi_{a}(\mathbf{2})\right| \frac{\mathbf{1}}{r_{12}}\left|\chi_{b}(\mathbf{2}) \chi_{a}(\mathbf{1})\right\rangle+\left\langle\chi_{a}(\mathbf{1}) \chi_{a}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{b}(\mathbf{2}) \chi_{b}(\mathbf{1})\right\rangle$
$+\left\langle\chi_{a}(\mathbf{1}) \chi_{b}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{a}(\mathbf{2}) \chi_{a}(\mathbf{1})\right\rangle+\left\langle\chi_{a}(\mathbf{1}) \chi_{b}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{a}(\mathbf{2}) \chi_{b}(\mathbf{1})\right\rangle$
$+\left\langle\chi_{a}(\mathbf{1}) \chi_{b}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{b}(\mathbf{2}) \chi_{a}(\mathbf{1})\right\rangle+\left\langle\chi_{a}(\mathbf{1}) \chi_{b}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{b}(\mathbf{2}) \chi_{b}(\mathbf{1})\right\rangle$
$+\left\langle\chi_{b}(\mathbf{1}) \chi_{a}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{a}(\mathbf{2}) \chi_{a}(\mathbf{1})\right\rangle+\left\langle\chi_{b}(\mathbf{1}) \chi_{a}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{a}(\mathbf{2}) \chi_{b}(\mathbf{1})\right\rangle$
$+\left\langle\chi_{b}(\mathbf{1}) \chi_{a}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{b}(\mathbf{2}) \chi_{a}(\mathbf{1})\right\rangle+\left\langle\chi_{b}(\mathbf{1}) \chi_{a}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{b}(\mathbf{2}) \chi_{b}(\mathbf{1})\right\rangle$
$+\left\langle\chi_{b}(\mathbf{1}) \chi_{b}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{a}(\mathbf{2}) \chi_{a}(\mathbf{1})\right\rangle+\left\langle\chi_{b}(\mathbf{1}) \chi_{b}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{a}(\mathbf{2}) \chi_{b}(\mathbf{1})\right\rangle$
$\left.+\left\langle\chi_{b}(1) \chi_{b}(2)\right| \frac{1}{r_{12}}\left|\chi_{b}(\mathbf{2}) \chi_{a}(1)\right\rangle+\left\langle\chi_{b}(\mathbf{1}) \chi_{b}(\mathbf{2})\right| \frac{1}{r_{12}}\left|\chi_{b}(2) \chi_{b}(1)\right\rangle\right]$

The 16 integrals above, $\boldsymbol{I}_{1-16}$, separate into four distinct types:
$\left\langle\chi_{a} \mathbf{( 1 )} \chi_{a}(\mathbf{2})\right| \frac{\mathbf{1}}{r_{12}}\left|\chi_{a}(\mathbf{2}) \chi_{a}(\mathbf{1})\right\rangle \quad$ electrons 1 and 2 both on a or both on $\mathrm{b}(\mathrm{x} 2) \quad\left(\boldsymbol{I}_{1}\right.$ and $\left.\boldsymbol{I}_{\mathbf{1 6}}\right)$ $\left\langle\chi_{a}(\mathbf{1}) \chi_{b}(\mathbf{2})\right| \frac{\mathbf{1}}{r_{12}}\left|\chi_{b}(\mathbf{2}) \chi_{a}(\mathbf{1})\right\rangle \quad$ electron 1 on a, electron 2 on b or vice versa $(\mathrm{x} 2)\left(\boldsymbol{I}_{7}, \boldsymbol{I}_{10}\right)$ $\left\langle\chi_{a}(\mathbf{1}) \chi_{b}(\mathbf{2})\right| \frac{\mathbf{1}}{r_{12}}\left|\chi_{a}(\mathbf{2}) \chi_{b}(\mathbf{1})\right\rangle \quad$ overlap densities of electrons 1 and $2(\mathrm{x} 4)\left(\boldsymbol{I}_{4,6,11,12}\right)$ $\left\langle\chi_{a}(\mathbf{1}) \chi_{a}(\mathbf{2})\right| \frac{\mathbf{1}}{r_{12}}\left|\chi_{a}(\mathbf{2}) \chi_{b}(\mathbf{1})\right\rangle \quad$ overlap density of electron 1 with electron 2 on a or on b , or vice versa (x8) ( $\left.\boldsymbol{I}_{2,3,5,8,9,13,14,15}\right)$
(relatively!) simple analytical solutions are available if you use a Gaussian basis set. Much harder solutions are available for Slater orbitals (see MQM P292), but very soon become intractable for higher angular momentum basis functions ( $p, d, f$ orbitals).

Comment: $\quad$ expanding $\Phi$ gives 4 terms
expanding $\hat{\boldsymbol{H}}$ splits each of these 4 into 3 components
( $2 \times 1$-electron $+1 \times 2$-electron), and the 1 -electron terms further split into 1 kinetic energy and two electron-nucleus integrals.
expanding $\sigma_{g}$ in terms of two atomic basis functions splits each of the each of the 1-e integrals into 4 components and each of the $2-\mathrm{e}$ integrals terms into 16 components. = a lot of integrals!

Note that in principle we can expand a molecular orbital using as many functions on as many atoms as we like:
$\varphi=N\left(\chi_{a}+\chi_{b}+\chi_{c}+\chi_{d}\right)$, for example (see 'basis sets')

If we do so, the number of integrals required increases rapidly (the number of 2-electron integrals scales as $\mathrm{N}^{4}$ where N is the number of basis functions). Moreover, the 1- and 2electron integrals ( $\boldsymbol{H}, \boldsymbol{J}$ and $\boldsymbol{K}$ ) can involve atomic functions on up to four different atoms, $\mathrm{a}, \mathrm{b}, \mathrm{c}$ and d :
e.g.

$$
\left\langle\chi_{a}(1) \chi_{b}(2)\right| \frac{1}{r_{12}}\left|\chi_{c}(1) \chi_{d}(2)\right\rangle=\langle i j \mid k l\rangle
$$

The 4-centre-2-electron integrals are numerous and very time-consuming to evaluate

## What have we achieved so far?

We have illustrated the machinery that allows the expression $\langle\Phi| \hat{\boldsymbol{H}}|\Phi\rangle=\boldsymbol{E}$ to be evaluated if we express the total wavefunction as an antisymmetrised product:

$$
\Phi=\left|1 \sigma_{g} 1 \bar{\sigma}_{g}\right|=\frac{1}{\sqrt{2}}\left|\begin{array}{ll}
1 \sigma_{g}(1) & 1 \bar{\sigma}_{g}(1) \\
1 \sigma_{g}(2) & 1 \bar{\sigma}_{g}(2)
\end{array}\right|
$$

and, critically, we know what the individual orbitals ( $1 \sigma_{g}$ in this case) look like. In the case of minimal basis set $\mathrm{H}_{2}$, symmetry is all we need ( $1 \sigma_{g}$ is the appropriately normalised in-phase combination of the 1s orbitals) but what if we don't know, a priori, what the orbitals look like. In $\mathrm{HHe}^{+}$, for example, we know that the occupied orbital is an in-phase combination of 1s orbitals and, qualitatively, we anticipate that the bonding orbital will be polarised towards the more electronegative He atom. But how polarised? 90\%? 60\%? The next section deals with Hartree-Fock theory, which provides a route, via the variation theorem, to find the optimum linear combination of atomic orbitals in cases where symmetry is not enough (i.e. the vast majority of problems!)

## Hartree-Fock theory

We have an expression for the total energy of a closed-shell molecule:

$$
E=\sum_{i} 2 E_{i}+\sum_{i, j}\left(2 J_{i j}-K_{i j}\right)
$$

where $\boldsymbol{E}_{\boldsymbol{i}}$ represents the interaction of an isolated electron with the nuclei. The 3-body problem ( 2 electrons + a fixed $\left(\mathrm{H}^{+}\right)_{2}$ unit) cannot be solved analytically because of the repulsion between the electrons.

Using the orbital approximation (i.e. assuming that each electron moves in an averaged potential defined by the nuclei and all other electrons - the 'mean field') we can express the ground-state wavefunction of $\mathrm{H}_{2}$ as a Slater determinant:
$\Psi=\left|1 \sigma_{g} 1 \bar{\sigma}_{g}\right|=\frac{1}{\sqrt{2}}\left|\begin{array}{ll}1 \sigma_{g}(\mathbf{1}) & 1 \bar{\sigma}_{g}(\mathbf{1}) \\ 1 \sigma_{g}(\mathbf{2}) & 1 \bar{\sigma}_{g}(2)\end{array}\right|$
But, in order to define the average field due to electron 1, you need to know its wavefunction (i.e. the form of the $1 \sigma_{g}$ orbital), and likewise for electron 2 . But what is the optimum form of the $1 \sigma_{g}$ orbital? (ans: we don't know yet, but certainly not the same as it was in $\mathrm{H}_{2}{ }^{+}$!).

We need to adopt an iterative approach - the 'self-consistent field'.

The formal derivation of the HF equations involves finding the condition under which $\boldsymbol{E}$, the total energy, is a minimum, subject to the constraint that the orbitals remain orthonormal, $\left\langle\phi_{i} \mid \phi_{j}\right\rangle=\delta_{i j}$ (the process involves the use of Lagrange multipliers - see box and MQM Ch7 further info for a full derivation).

## Lagrange multipliers: Finding an extremum subject to a constraint:

Find the turning point of $f=3 \boldsymbol{x}^{2}-\mathbf{2} \boldsymbol{y}^{2}$ subject to the constraint that $\boldsymbol{x}+\boldsymbol{y}=\mathbf{2}$

Soln: define $g=x+y$ and $\lambda$ so $f-\lambda g=3 x^{2}-2 y^{2}-\lambda(x+y)$
$\lambda$ is the 'Lagrange multiplier'

$$
\frac{\partial(f-\lambda g)}{d x}=6 x-\lambda=0 \quad \frac{\partial(f-\lambda g)}{d y}=-4 y-\lambda=0
$$

Ans: $\quad x=-4, y=6, \lambda=-24 \quad$ n.b. $f(-4,6)=3 x^{2}-2 y^{2}=-24=\lambda$

When we do this, we arrive at the Hartree-Fock equations,

$$
\hat{\boldsymbol{F}} \phi_{i}=\varepsilon_{i} \phi_{i}
$$

where the Fock operator, $\hat{\boldsymbol{F}}_{i}$, is defined as $\hat{\boldsymbol{F}}=\hat{\boldsymbol{H}}+\sum_{j} \mathbf{2} \hat{\boldsymbol{J}}_{j}-\sum_{j} \hat{\boldsymbol{K}}_{j}=\hat{\boldsymbol{H}}+\hat{\boldsymbol{G}}$
$\varepsilon_{i}$ is the Hartree Fock orbital energy (the Lagrange multiplier and also the eigenvalue)
$\hat{\boldsymbol{H}}$ is the one-electron term $\hat{\boldsymbol{H}}=-\frac{\mathbf{1}}{\mathbf{2}} \nabla_{1}{ }^{2}-\frac{\mathbf{1}}{\boldsymbol{r}_{a 1}}-\frac{\mathbf{1}}{\boldsymbol{r}_{b 1}}$
$\hat{\boldsymbol{J}}_{\boldsymbol{j}}$ and $\hat{\boldsymbol{K}}_{\boldsymbol{j}}$ are the Coulomb and Exchange operators, respectively

$$
\hat{J}_{j} \phi_{i}(1)=\left(\int \phi_{j}(2) \phi_{j}(2) \frac{1}{r_{12}} d v_{2}\right) \phi_{i}(1) \quad \hat{K}_{j} \phi_{i}(1)=\left(\int \phi_{j}(2) \phi_{i}(2) \frac{1}{r_{12}} d v_{2}\right) \phi_{j}(1)
$$

The Coulomb operator defines the influence of the charge cloud $\phi_{j}(\mathbf{2}) \phi_{j}(\mathbf{2})$ on electron I etc. The exchange operator defines the modification of this repulsion by spin correlation.

Thus $\varepsilon_{i}=\boldsymbol{E}_{i}+\sum_{j}\left(\mathbf{2 J} \boldsymbol{J}_{i j}-\boldsymbol{K}_{i j}\right)$

## Note the difference between $E_{i}$ and $\varepsilon_{i}$ :

$\boldsymbol{E}_{\boldsymbol{i}}$ is the interaction between the electron and the nuclei in the absence of the other electron (i.e. exactly as in $\mathrm{H}_{2}{ }^{+}$). $\varepsilon_{i}$ is the Hartree Fock orbital energy, which represents the energy of the electron in the combined field of the nuclei and the other electron.

We now have a system of linear equations where the Fock operator (a one-electron operator) defines the average field due to the nuclei (contained in $\hat{\boldsymbol{H}}$ ) and the remaining electrons (contained in $\hat{\boldsymbol{J}}$ and $\hat{\boldsymbol{K}}$ operators).

What do the one-electron energies, $\varepsilon_{i}$, mean?
Each orbital energy contains the effects of repulsions by all other electrons, so the sum over all occupied orbital energies, $\sum_{i} \varepsilon_{i}$, double counts the repulsions. Therefore, to calculate the total energy we have to correct for this by subtracting the e-e repulsion.
$\boldsymbol{E}=\sum_{i} \mathbf{2} \varepsilon_{i}-\sum_{i, j}\left(\mathbf{2} J_{i j}-\boldsymbol{K}_{i j}\right)$ where the sum is over doubly-occupied orbitals.
Check this works for $\mathrm{H}_{2}$ :
$E=2 \varepsilon_{1 \sigma_{g}}-2 J_{1 \sigma_{g} 1 \sigma_{g}}+K_{1 \sigma_{g} 1 \sigma_{g}}$
but recall that $\quad \boldsymbol{J}_{1 \sigma_{g} 1 \sigma_{g}}=\boldsymbol{K}_{1 \sigma_{g} 1 \sigma_{g}}$
and that

$$
\begin{aligned}
& \varepsilon_{1 \sigma_{g}}=E_{1 \sigma_{g}}+J_{1 \sigma_{g} 1 \sigma_{g}} \\
& \therefore \boldsymbol{E}=\mathbf{2}\left(\boldsymbol{E}_{1 \sigma_{g}}+J_{1 \sigma_{g} 1 \sigma_{g}}\right)-\mathbf{2} J_{1 \sigma_{g} 1 \sigma_{g}}+J_{1 \sigma_{g} 1 \sigma_{g}}=2 E_{1 \sigma_{g}}+J_{1 \sigma_{g} 1 \sigma_{g}} \text { (as previous). }
\end{aligned}
$$

Koopmans' theorem: "The ionisation energy is the negative of the Hartree Fock orbital energy": $\quad \boldsymbol{I}_{i}=-\varepsilon_{i}$

Why?

$$
I_{H_{2}}=E\left(H_{2}^{+}\right)-E\left(H_{2}\right)=\left(E_{1 \sigma_{g}}\right)-\left(2 E_{1 \sigma_{g}}+J_{1 \sigma_{g} 1 \sigma_{g}}\right)=-\left(E_{1 \sigma_{g}}+J_{1 \sigma_{g} 1 \sigma_{g}}\right)=-\varepsilon_{1 \sigma_{g}}
$$

Another way to see this is to note that when we remove an electron we lose the attraction to the nuclei ( $\boldsymbol{E}_{1 \sigma_{g}}$ ) and also 1 unit of Coulomb repulsion - precisely the components of $\boldsymbol{\varepsilon}_{1 \sigma_{g}}$

What do the energies of the vacant orbitals mean? Consider the $\sigma_{u}$ orbital in $\mathrm{H}_{2}$.

$$
\varepsilon_{1 \sigma_{u}}=E_{1 \sigma_{u}}+\sum_{i}\left(2 J_{i j}-K_{i j}\right)=E_{1 \sigma_{u}}+2 J_{1 \sigma_{g} 1 \sigma_{u}}-K_{1 \sigma_{g} 1 \sigma_{u}}
$$

i.e. it experiences repulsions from both electrons (unlike $\sigma_{g}$ which only experiences repulsion from one electron (the other being 'itself'). So beware - it is dangerous to compare occupied and virtual HF orbital energies.

The energy of a virtual orbital in HF theory corresponds to the negative of the electron affinity of the atom/molecule.

$$
\begin{aligned}
& \boldsymbol{E} A_{H_{2}}=\boldsymbol{E}\left(\boldsymbol{H}_{2}\right)-\boldsymbol{E}\left(\boldsymbol{H}_{2}^{-}\right)=\left(2 E_{1 \sigma_{g}}+\boldsymbol{J}_{1 \sigma_{g} 1 \sigma_{g}}\right)-\left(2 E_{1 \sigma_{g}}+\boldsymbol{E}_{1 \sigma_{u}}+J_{1 \sigma_{g} 1 \sigma_{g}}+2 J_{1 \sigma_{g} 1 \sigma_{u}}-\boldsymbol{K}_{1 \sigma_{g} 1 \sigma_{u}}\right) \\
& =-\left(\boldsymbol{E}_{1 \sigma_{u}}+2 J_{1 \sigma_{g} 1 \sigma_{u}}-\boldsymbol{K}_{1 \sigma_{g} 1 \sigma_{u}}\right)=-\varepsilon_{1 \sigma_{u}}
\end{aligned}
$$

## Hartree-Fock-Roothaan (HFR) Equations:

These are the HF equations implemented using the LCAO approximation $\varphi=\sum_{n} c_{n} \chi_{n}$

$$
\hat{F} \sum_{n} c_{n} \chi_{n}=\varepsilon_{i} \sum_{n} c_{n} \chi_{n}
$$

Multiply by $\chi_{m}$ and integrating gives $\sum_{n} c_{n}\left\langle\chi_{m}\right| \hat{\boldsymbol{F}}\left|\chi_{n}\right\rangle=\varepsilon_{i} \sum_{n}\left\langle\chi_{m}\right| c_{n}\left|\chi_{n}\right\rangle$
$\sum_{n} c_{n}\left(\boldsymbol{F}_{m n}-\varepsilon_{i} S_{m n}\right)=\mathbf{0}$ which have non-trivial solutions where $\left|\boldsymbol{F}_{m n}-\varepsilon_{i} \boldsymbol{S}_{m n}\right|=\mathbf{0}$
(note similarities to secular equations derived in Lecture 1 - more on solving these in the last 2 lectures).

Reiterate: in order to determine $c_{n}$ by solving the HFR equations, we need to be able to calculate the matrix elements $\boldsymbol{F}_{m n}$ (including Coulomb and exchange integrals), which in turn means that we need to already know $\boldsymbol{c}_{\boldsymbol{n}}$ ! Hence an iterative solution required:

In the context of HF theory, the flow diagram is as follows:


## An example: $\mathrm{H}_{2}$ at the equilibrium separation, $\mathrm{R}=1.38543 \mathrm{a}_{0}$.

Expand the MOs in the form $\phi=c_{a} \chi_{a}+c_{b} \chi_{b}$ where $\chi_{a}, \chi_{b}$ are 1 s orbitals centred on the H atoms. We will use a normalised 'Slater-type' orbital of the form $\chi_{1 s}=N e^{-\zeta r}$ (see 'basis functions') with exponent $\zeta=\mathbf{1 . 1 9 3 0 2}$ (chosen because it is the optimum value at this separation, but the choice doesn't matter for our purposes here).
The general expression for the HFR equations $\sum_{n} c_{n}\left(F_{m n}-\varepsilon_{i} S_{m n}\right)=0$ give us

$$
\begin{aligned}
& \left(F_{a a}-\varepsilon S_{a a}\right) c_{a}+\left(F_{a b}-\varepsilon S_{a b}\right) c_{b}=0 \\
& \left(F_{a b}-\varepsilon S_{a b}\right) c_{a}+\left(F_{b b}-\varepsilon S_{b b}\right) c_{b}=0
\end{aligned} \quad \text { and } \quad\left|\begin{array}{ll}
F_{a a}-\varepsilon S_{a a} & F_{a b}-\varepsilon S_{a b} \\
F_{a b}-\varepsilon S_{a b} & F_{b b}-\varepsilon S_{b b}
\end{array}\right|=0
$$

numerical values for matrix elements ( $\boldsymbol{S}, \boldsymbol{H}, \boldsymbol{G}$ in Hartree units) as function of $\boldsymbol{c}_{a, b}$

$$
\begin{gathered}
S_{a a}=S_{b b}=1 \\
S_{a b}=\mathbf{0 . 6 8 2 4 2} \\
\boldsymbol{F}_{a a}=\boldsymbol{H}_{a a}+\boldsymbol{G}_{a a}=\left(-\mathbf{1 . 1 3 2 9 5}+\mathbf{0 . 7 4 5 6 4} c_{a}{ }^{2}+\mathbf{0 . 8 9 3 5 9} c_{a} c_{b}+\mathbf{0 . 8 1 3 3 5} c_{b}{ }^{2}\right) a u \\
\boldsymbol{F}_{b b}=\boldsymbol{H}_{b b}+\boldsymbol{G}_{b b}=\left(-\mathbf{1 . 1 3 2 9 5}+\mathbf{0 . 8 1 3 3 5} c_{a}{ }^{2}+\mathbf{0 . 8 9 3 5 9} c_{a} c_{b}+\mathbf{0 . 7 4 5 6 4} c_{b}{ }^{2}\right) a u \\
\boldsymbol{F}_{a b}=\boldsymbol{H}_{a b}+\boldsymbol{G}_{a b}=\left(-\mathbf{0 . 9 7 4 7 5}+\mathbf{0 . 4 4 6 7 5} c_{a}{ }^{2}+\mathbf{0 . 3 6 4 9 8} c_{a} c_{b}+\mathbf{0 . 4 4 6 7 5} c_{b}{ }^{2}\right) a u
\end{gathered}
$$

(the numbers come from well-established formulae: the details are not important but are contained in Szabo and Ostlund)

All we need to start the iteration is a guess for $c_{a}$ and $c_{b}$. Clearly in this case we could guess from symmetry considerations that $c_{a}=c_{b}$ for the ground state, but to illustrate the process, let's start with an obviously incorrect guess:

Step 1: guess

$$
c_{a}=2 c_{b} \quad \therefore \frac{c_{a}}{c_{b}}=2
$$

Step 2: normalise $\quad \phi=2 c_{b} \chi_{a}+c_{b} \chi_{b}$

$$
\begin{aligned}
& \int\left(2 c_{b} \chi_{a}+c_{b} \chi_{b}\right)\left(2 c_{b} \chi_{a}+c_{b} \chi_{b}\right) d \tau=1 \\
& \Rightarrow 5 c_{b}^{2}+4 c_{b}^{2} S_{a b}=1 \\
& \Rightarrow c_{b}=\frac{1}{\sqrt{4 S_{a b}+5}}=0.35968 \\
& \Rightarrow c_{a}=0.71936
\end{aligned}
$$

Step 3: calculate matrix elements using formulae for $\boldsymbol{F}_{a a}, \boldsymbol{F}_{a b}, \boldsymbol{F}_{b b}$ above

$$
\begin{aligned}
& F_{a a}=H_{a a}+G_{a a}=\left(-1.13295+0.74564(0.71936)^{2}+0.89359(0.71936)(0.35968)+0.81335(0.35968)^{2}\right) a u \\
& =-0.4107 a u
\end{aligned}
$$

$$
\left|\begin{array}{cc}
-0.4107-\varepsilon & -0.5913-0.6824 \varepsilon \\
-0.5913-0.6824 \varepsilon & -0.3844-\varepsilon
\end{array}\right|=0
$$

Step 4: solve $\quad \varepsilon_{1}=\mathbf{- 0 . 5 8 8 0 4 a u}$ (lowest eigenvalue)

Step 5: Converged? (i.e. is the value of $\varepsilon_{1}$ the same (to within a defined threshold) as the previous value). Clearly as this is the first iteration we have nothing to compare to, and so the answer is no!

Otherwise
Step 6: solve

$$
\left.\begin{array}{lr}
\left(F_{a a}-\varepsilon S_{a a}\right) c_{a}+\left(F_{a b}-\varepsilon S_{a b}\right) c_{b}=0 \\
\left(F_{a b}-\varepsilon S_{a b}\right) c_{a}+\left(F_{b b}-\varepsilon S_{b b}\right) c_{b}=0
\end{array}\right\} \text { for } c_{a}, c_{b} \text { with } \varepsilon=\varepsilon_{1}=-\mathbf{0 . 5 8 8 0 4 a u} \text {. } \begin{array}{lr}
\mathbf{0 . 1 7 7 3} c_{a}-\mathbf{0 . 1 9 9 0} c_{b}=0 & \text { Convergence history } \\
\begin{array}{ll}
c_{a} \\
c_{b} & =1.0715
\end{array}
\end{array}
$$

go round the cycle again...........

Step 2: normalise
$c_{a}{ }^{2}+2 c_{a} c_{b} S+c_{b}{ }^{2}=1 \Rightarrow$
$c_{a}=0.56391$
$c_{b}=0.52627$

|  | $\boldsymbol{c}_{\boldsymbol{a}}$ | $\boldsymbol{c}_{\boldsymbol{a}}$ | $\boldsymbol{F}_{\boldsymbol{a} \boldsymbol{a}}$ | $\boldsymbol{F}_{\boldsymbol{a b}}$ | $\boldsymbol{F}_{\boldsymbol{b} \boldsymbol{b}}$ | $\boldsymbol{\varepsilon}_{\mathbf{1}}$ | $\boldsymbol{\varepsilon}_{\mathbf{2}}$ |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| Initial | 0.71937 | 0.35968 | -0.4107 | -0.5913 | -0.3844 |  |  |
| $\mathbf{1}^{\text {st }}$ iter | 0.56391 | 0.52627 | -0.4054 | -0.6001 | -0.4026 | -0.5880 | 0.6105 |
| $\mathbf{2}^{\text {nd }}$ iter | 0.54711 | 0.54319 | -0.4042 | -0.6007 | -0.4039 | -0.5972 | 0.6191 |
| $\mathbf{3}^{\text {rd }}$ iter | 0.54536 | 0.54495 | -0.4041 | -0.6007 | -0.4041 | -0.5973 | 0.6192 |
| $\mathbf{4}^{\text {th }}$ iter | 0.54517 | 0.54513 | -0.4041 | -0.6007 | -0.4041 | -0.5973 | 0.6192 |
| $\mathbf{5}^{\text {th }}$ iter | 0.54515 | 0.54515 | -0.4041 | -0.6007 | -0.4041 | -0.5973 | 0.6192 |

etc. etc. etc. etc.

## The concept of effective nuclear charge: the 4s vs 3d controversy.

For a poly-electronic electron, the full Hamiltonian for a single electron is:

$$
H_{e l, i}=-\frac{1}{2} \nabla^{2}-\frac{Z}{r_{i}}+\sum_{j>i}^{n} \frac{1}{r_{i j}}
$$

But we can simplify this by absorbing the (repulsive) electron-electron interactions into the 'effective nuclear charge", which is reduced from its value for the real atom:

$$
H_{e l, i} \approx-\frac{1}{2} \nabla^{2}-\frac{Z^{e f f}}{r_{i}}
$$

We say that the reference electron (i) is 'shielded' from the real nuclear charge $Z$ by the presence of the other electrons, and if we choose an appropriate value of $Z^{\text {eff }}$, we can recover the eigenvalues and eigenfunctions of the full Hamiltonian. There are many sets of empirical rules used to define $Z^{\text {eff }}$ (Slater, Clementi etc).

So we really 'hide' the electron-electron repulsions by rolling then into a reduced positive charge at the nucleus: the 'effective nuclear charge'. This was absolutely necessary when the concept was first thought of (ca 1930), because accurate computation of the repulsions was out of the question. Now they are routine.

Qu 1: The $4 s / 3 d$ ordering debate: why is $K 4 s^{1} 3 d^{0}$ and not $3 d^{1} 4 s^{0}$ ?
Standard answer: the 4 s orbital penetrates through the core electrons, and so 'experiences a greater effective nuclear charge'.

Qu 2: why is $T i^{3+} 3 d^{1} 4 s^{0}$ and not $4 s^{1} 3 d^{0}$ ?
Standard answer: don't know.

Let us consider the ${ }^{7} \mathrm{~S}$ states of the isoelectronic series $\mathrm{Cr}, \mathrm{Mn}^{+}, \mathrm{Fe}^{2+}$, where the 4 s and 3 d orbitals are all singly occupied (thereby avoiding the problem of comparing occupied and virtual orbital energies).

Components of the self-consistent energies of the 4 s and 3 d orbitals of the ${ }^{7} \mathbf{S}$ states $\mathbf{o f} \mathbf{C r}$, $\mathrm{Mn}^{+}$and $\mathrm{Fe}^{2+}$ (in au) at the HF/VTZ level.
$\boldsymbol{T} \quad V_{n e} \quad V_{e e} \quad E_{t o t}$

Cr

| 4s | 0.63 | -8.29 | 7.44 | -0.22 |
| :---: | :---: | :---: | :---: | :---: |
| 3d | 4.91 | -24.13 | 18.86 | -0.36 |

$\mathbf{M n}^{+}$

| 4s | 1.13 | -10.96 | 9.28 | -0.54 |
| :--- | :---: | :---: | :---: | :---: |
| 3d | 6.50 | -29.26 | 21.84 | -0.93 |

$\mathrm{Fe}^{2+}$

| 4s | 1.71 | -13.52 | 10.86 | -0.95 |
| :--- | :--- | :--- | :--- | :--- |
| 3d | 8.13 | -34.25 | 24.47 | -1.65 |

$$
\operatorname{Cr}\left(4 s^{1} 3 d^{5}\right)
$$



Components of the electron-electron repulsion in the ${ }^{7} \mathrm{~S}$ state of Cr .

The correlation problem and post-HF methods.
We have established from MO theory that the ground state wavefunction for $\mathrm{H}_{2}$ can be represented as:

$$
\Phi_{M O}=\left|1 \sigma_{g} 1 \bar{\sigma}_{g}\right|=\frac{1}{\sqrt{2}}\left|\begin{array}{ll}
1 \sigma_{g}(1) & 1 \bar{\sigma}_{g}(1) \\
1 \sigma_{g}(2) & 1 \bar{\sigma}_{g}(2)
\end{array}\right|
$$

and we used the HFR equations to find the optimum form of the orbital $1 \sigma_{g}$ in terms of the LCAO expansion. The potential energy surface shown below ('HF' = Hartree Fock) indicates that this reproduces the equilibrium geometry reasonably well, but fails to reproduce the dissociation energy: the HF energy of $\mathrm{H}_{2}$ at the dissociation limit is $\sim 0.34$ au higher than that of two isolated H atoms (set as zero in the figure below).


To see the origin of the problem, expand the spatial part of the wavefunction:

$$
\begin{aligned}
& \Phi_{\text {spatial }}=\mathbf{1} \sigma_{g} 1 \sigma_{g}=\frac{1}{2(1+S)}\left(\chi_{a}(1)+\chi_{b}(1)\right)\left(\chi_{a}(\mathbf{2})+\chi_{b}(\mathbf{2})\right) \\
& =\frac{1}{2(1+S)}\left(\chi_{a}(1) \chi_{a}(2)+\chi_{b}(1) \chi_{b}(2)+\chi_{a}(1) \chi_{b}(2)+\chi_{b}(1) \chi_{a}(2)\right)
\end{aligned}
$$

i.e. an equal mixture of covalent $\left(H_{a}{ }^{*} \ldots . \mathbf{H}_{b}{ }^{*}\right)$ and ionic $\left(\mathbf{H}_{a}{ }^{+} \ldots . \mathbf{H}_{b}{ }^{-}\right)$resonance structures. This is reasonable at the equilibrium geometry, where overlap is large, but unrealistic at the dissociation limit, where two neutral H atoms will be strongly favoured over $\mathbf{H}^{+}+\mathbf{H}^{-}$.

This problem is a consequence of the central field approximation, where one electron is influenced only by the average position of the other: the result is that situations where both electrons are very close to each other at any given instant are more probable than they should be: we say that the motion of the two electrons is not correlated.
Another way to see this is to note that if the two electrons share the same spatial wavefunction, then wherever the most probable position to find electron 1 is, it is also the most probable position to find electron 2, and that doesn't make much sense!

## Configuration interaction

The problems at the dissociation limit can be reduced by using configuration interaction adding in extra configurations to the wavefunction.
In this case we will mix the $\left(1 \sigma_{g}\right)^{2}$ configuration with the doubly excited configuration $\left(1 \sigma_{u}\right)^{2}$

$$
\Psi_{C I}=c_{1}\left|1 \sigma_{g} 1 \bar{\sigma}_{g}\right|+c_{2}\left|1 \sigma_{u} 1 \bar{\sigma}_{u}\right|
$$

Expanding the spatial part of the second term in the same way as above gives

$$
\begin{aligned}
& \Phi_{\text {spatial }}=1 \sigma_{u} 1 \sigma_{u}=\frac{1}{2(1-S)}\left(\chi_{a}(1)-\chi_{b}(1)\right)\left(\chi_{a}(2)-\chi_{b}(2)\right) \\
& =\frac{1}{2(1-S)}\left(\chi_{a}(1) \chi_{a}(2)+\chi_{b}(1) \chi_{b}(2)-\chi_{a}(1) \chi_{b}(2)-\chi_{b}(1) \chi_{a}(2)\right)
\end{aligned}
$$

$$
\begin{aligned}
& \text { Thus } \Psi_{C I}=c_{1}\left|1 \sigma_{g} 1 \bar{\sigma}_{g}\right|+c_{2}\left|1 \sigma_{u} 1 \bar{\sigma}_{u}\right| \\
& \Psi_{\text {spatial }}=\left(\frac{c_{1}}{2(1+S)}+\frac{c_{2}}{2(1-S)}\right)\left(\chi_{a}(1) \chi_{a}(2)+\chi_{b}(1) \chi_{b}(2)\right) \\
& +\left(\frac{c_{1}}{2(1+S)}-\frac{c_{2}}{2(1-S)}\right)\left(\chi_{a}(1) \chi_{b}(2)+\chi_{b}(1) \chi_{a}(2)\right)
\end{aligned}
$$

Thus an appropriate choice of coefficients (specifically when $\frac{c_{1}}{c_{2}}=-\frac{(1+S)}{(1-S)}$ ) causes the ionic terms to vanish completely. The ratio $\frac{c_{2}}{c_{1}}$ varies from 0 at equilibrium to -1 at the dissociation limit.
The correlation energy is defined as the difference between the energy obtained from this wavefunction and the Hartree Fock energy. Notice that in order to get a better energy, we have had to abandon the idea that electrons live in orbitals - this is a big leap!

In practice, Cl is performed by constructing linear combinations of ground and excited determinants based on the Hartree-Fock orbitals. The optimum linear combination

$$
\Psi=\sum_{i} C_{i} \Phi_{i}
$$

where $\Phi_{i}$ are now determinants, not orbitals, is then found using the linear variation method.

Example: for $\mathrm{H}_{2}: \Psi=\boldsymbol{C}_{1} \Phi_{1}+\boldsymbol{C}_{2} \Phi_{2}$

$$
\Phi_{1}=\left|1 \sigma_{g} 1 \bar{\sigma}_{g}\right|=\frac{1}{\sqrt{2}}\left|\begin{array}{ll}
1 \sigma_{g}(1) & 1 \bar{\sigma}_{g}(1) \\
1 \sigma_{g}(2) & 1 \bar{\sigma}_{g}(2)
\end{array}\right| \quad \Phi_{2}=\left|1 \sigma_{u} 1 \bar{\sigma}_{u}\right|=\frac{1}{\sqrt{2}}\left|\begin{array}{ll}
1 \sigma_{u}(1) & 1 \bar{\sigma}_{u}(1) \\
1 \sigma_{u}(2) & \mathbf{1} \bar{\sigma}_{u}(2)
\end{array}\right|
$$

and the secular determinant we need to solve is:

$$
\left\lvert\, \begin{array}{cc}
\left\langle\Phi_{1}\right| \hat{\boldsymbol{H}}\left|\Phi_{1}\right\rangle-\boldsymbol{E} & \left\langle\Phi_{2}\right| \hat{\boldsymbol{H}}\left|\Phi_{1}\right\rangle \\
\left\langle\Phi_{1}\right| \hat{\boldsymbol{H}}\left|\Phi_{2}\right\rangle & \left\langle\Phi_{2}\right| \hat{\boldsymbol{H}}\left|\Phi_{2}\right\rangle-\boldsymbol{E} \mid
\end{array}=\mathbf{0}\right.
$$

We have already dealt with the matrix element $\left\langle\Phi_{1}\right| \boldsymbol{H}\left|\Phi_{1}\right\rangle$ - it is the ground-state energy of $\mathrm{H}_{2}:=\mathbf{2} \boldsymbol{E}_{1 \sigma_{g}}+\boldsymbol{J}_{1 \sigma_{g} 1 \sigma_{g}}$
Similarly, we can show that

$$
\left\langle\Phi_{2}\right| \hat{\boldsymbol{H}}\left|\Phi_{2}\right\rangle=\mathbf{2} \boldsymbol{E}_{1 \sigma_{u}}+\boldsymbol{J}_{1 \sigma_{u} 1 \sigma_{u}} \text { and }\left\langle\Phi_{1}\right| \hat{\boldsymbol{H}}\left|\Phi_{2}\right\rangle=\boldsymbol{K}_{1 \sigma_{g} 1 \sigma_{u}} \text { (see problems sheet) }
$$

So the total energy including Cl is the lowest root of

$$
\left|\begin{array}{cc}
2 E_{1 \sigma_{g}}+J_{1 \sigma_{g} 1 \sigma_{g}}-\boldsymbol{E} & K_{1 \sigma_{g} 1 \sigma_{u}} \\
K_{1 \sigma_{g} 1 \sigma_{u}} & 2 E_{1 \sigma_{u}}+J_{1 \sigma_{u} 1 \sigma_{u}}-\boldsymbol{E}
\end{array}\right|=0
$$

This is the dashed curve in the Figure on P 29

For larger molecules, many millions or even billions of excited determinants may be needed for good accuracy. The methods (acronyms CI, CID, CISD(T), CASSCF...) can be very accurate but very expensive.


## Density Functional Theory

A fundamentally different approach, based on the density, $\rho$, (a 3-dimensional variable) rather than the wavefunction (a $3 n$-dimensional entity, where $n$ is the number of electrons). Based on Theorems by Hohenberg and Kohn (Kohn, Nobel prize 1998):
a) The energy is uniquely defined by the electron density, $\rho$.
b) The ground-state energy can be obtained variationally: the density that minimises the total energy is the exact ground-state density.

In principle, orbitals are not needed for DFT, but in practice it is very difficult to calculate the KE term ( $-\frac{\mathbf{1}}{\mathbf{2}} \nabla^{2}$ ) accurately using the density alone. Kohn and Sham developed a methodology wherein the electron density is modelled as a system of fictitious noninteracting electrons which occupy the 'Kohn-Sham orbitals') and reproduce the real density. This leads to a set of 1-electron Kohn-Sham equations very similar in structure to the Hartree Fock equations:
$\hat{\boldsymbol{H}}_{\text {KS }} \phi_{i}=\left(-\frac{\mathbf{1}}{\mathbf{2}} \nabla^{2}+\boldsymbol{V}_{\text {eff }}\right) \phi_{i}=\varepsilon_{i} \phi_{i} \quad \boldsymbol{V}_{\text {eff }}=\boldsymbol{V}(\boldsymbol{r})+\boldsymbol{J}(\boldsymbol{r})+\boldsymbol{V}_{x c}(\boldsymbol{r})$
$V(r), \boldsymbol{J}(r)$ are the electron-nuclear and coulomb potentials, precisely as in HF theory. $\boldsymbol{V}_{x c}(\boldsymbol{r})$ is the so-called 'exchange-correlation' potential, which determines both the exchange and correlation effects.

Notice that the influence of correlation is introduced through the Hamiltonian, not by expanding the wavefunction beyond the HF single determinant. Therefore it is appealing in terms of cost.

Problem: $\quad V_{x c}(r)$ is unknown, and there are hundreds (probably thousands now) of different approximations to it (BP86, BLYP, HCTH, B3LYP, PBE, rev-PBE........ - the socalled 'functional zoo'). More than 20 years after it was first proposed, the most popular remains B3LYP ('everybody's favourite functional').

## Semi-empirical theory

The number of 2-electron integrals in a HF expansion scales as $N^{4}$ where $N$ is the number of basis functions. Thus reducing the size of the basis set and avoiding the calculation of some of the integrals is clearly an advantageous strategy.

## Hückel theory (see Valence lectures 7/8)

The ultimate extension of semi-empirical theory - the philosophy is to eliminate the computation of matrix elements entirely by parameterisation at the Fock matrix element level. Thus it is not iterative.

1) Assume complete separation of $\sigma / \pi$ systems (contrast EHT, later, where all valence orbitals are included)
2) Assumes independent electrons: i.e. $\Phi=\varphi_{1} \varphi_{2} \varphi_{3} \varphi_{4} \ldots .$. , a simple Hartree product, and $\boldsymbol{E}=\boldsymbol{E}_{1}+\boldsymbol{E}_{2}+\boldsymbol{E}_{3}+\boldsymbol{E}_{4}+\ldots$.
3) 
4) $\quad \boldsymbol{H}_{12}=\boldsymbol{\beta} \quad$ (the resonance integral) if 1 and 2 are nearest neighbours, 0 otherwise
n.b $\alpha$ and $\beta$ usually given symbolic values rather than assigned to real numbers.
5) $\quad S_{12}=0$
this makes the overlap matrix equal to the identity, and simplifies the secular equations from $\hat{\boldsymbol{H}} \boldsymbol{c}=\boldsymbol{S} \boldsymbol{c} \boldsymbol{E}$ to $\hat{\boldsymbol{H}} \boldsymbol{c}=\boldsymbol{c} \boldsymbol{E}$. In a typical $\pi$ system $\boldsymbol{S}_{12} \approx \mathbf{0 . 2}$, so this seems a radical approximation. However it has little impact on the final solutions (see problem 2.2).

Note geometry (i.e. bond lengths) is not considered at any point - only connectivity.

Example: allyl radical

$$
\left|\begin{array}{ccc}
\alpha-E & \beta & 0 \\
\beta & \alpha-E & \beta \\
0 & \beta & \alpha-E
\end{array}\right|=0
$$

Define $x=\left(\frac{\alpha-E}{\beta}\right) \Rightarrow\left|\begin{array}{ccc}x & 1 & 0 \\ 1 & x & 1 \\ 0 & 1 & x\end{array}\right|=x^{3}-2 x=0$

$$
x=0, \pm \sqrt{2} \quad x=0 \Rightarrow E_{2}=\alpha \quad x= \pm \sqrt{2} \Rightarrow E_{1,3}=\alpha \pm \sqrt{2} \beta
$$

note energy level subscripts 1-3 chosen for future convenience.

Eigenvectors: sub $\boldsymbol{E}_{1-3}$ into secular equations
$E_{2}=\boldsymbol{\alpha}$
$\left.\begin{array}{l}c_{1} \beta+c_{2}(\alpha-\alpha)+c_{3} \beta=0 \Rightarrow c_{3}=-c_{1} \\ c_{1}(\alpha-\alpha)+c_{2} \beta=0 \Rightarrow c_{2}=0\end{array}\right\} \quad \psi_{2}=\frac{1}{\sqrt{2}}\left(\chi_{1}-\chi_{3}\right) \quad$ (after normalisation)
$\boldsymbol{E}_{1}=\alpha+\sqrt{2} \beta$
$\left.\begin{array}{l}c_{1}(\alpha-\alpha-\sqrt{2} \beta)+c_{2} \beta=0 \Rightarrow c_{2}=\sqrt{2} c_{1} \\ c_{1} \beta+\sqrt{2} c_{1}(\alpha-\alpha-\sqrt{2} \beta)+c_{3} \beta=0 \Rightarrow c_{3}=c_{1}\end{array}\right\} \quad \psi_{1}=\frac{1}{2}\left(\chi_{1}+\sqrt{2} \chi_{2}+\chi_{3}\right)$
$\boldsymbol{E}_{3}=\alpha-\sqrt{2} \beta$
$\left.\begin{array}{l}c_{1}(\alpha-\alpha+\sqrt{2} \beta)+c_{2} \beta=0 \Rightarrow c_{2}=-\sqrt{2} c_{1} \\ c_{1} \beta-\sqrt{2} c_{1}(\alpha-\alpha+\sqrt{2} \beta)+c_{3} \beta=0 \Rightarrow c_{3}=c_{1}\end{array}\right\} \quad \psi_{3}=\frac{1}{2}\left(\chi_{1}-\sqrt{2} \chi_{2}+\chi_{3}\right)$

$\psi_{2}$

$\psi_{1}$

Singly occupied orbital is $E_{2}$ : the unpaired electron lies exclusively on the outer carbons.

## Extended Hückel theory

As with Hückel theory, the computation of matrix elements is avoided entirely by parameterisation of the Fock matrix elements: again not iterative.
But no assumptions are made about $\sigma / \pi$ separability: all valence electrons included. Unlike Hückel theory, geometry determines overlaps, which are calculated explicitly (clearly ignoring $\sigma$ overlap is not reasonable!).
Basis of method

1) Assumes independent electrons: $\Phi=\varphi_{1} \varphi_{2} \varphi_{3} \varphi_{4} \ldots$. and $E=E_{1}+E_{2}+E_{3}+E_{4}+\ldots$.
2) $\quad S_{i j}$ computed as in Hartree Fock theory, using a minimal basis of Slater-type orbitals. If you want to see the details of how this is done, look at the python script EH.py in the course support materials.
3) $\quad \boldsymbol{H}_{i i}=-\boldsymbol{I}_{i}$ The ionisation energy of an electron in the appropriate orbital in the valence configuration (e.g. $\mathrm{sp}^{3}$ for a tetrahedral carbon). In practice this is usually taken as the weighted average of $2 s$ and $2 p$ electrons in $C$, but the choice is not critical.
4) $\quad \boldsymbol{H}_{i j}=\boldsymbol{K} \boldsymbol{S}_{i j} \frac{\left(\boldsymbol{H}_{i i}+\boldsymbol{H}_{i j}\right)}{2} \quad \boldsymbol{K}=\mathbf{1 . 7 5} \quad$ Wolfsberg-Helmholtz formula (1.75 reproduces rotational barrier for ethane, but again choice not critical)

Justification: interaction will increase with overlap, and with the energies of the component orbitals (for a fixed overlap).
n.b. early versions used the simpler $\boldsymbol{H}_{i j}=\boldsymbol{K} \boldsymbol{S}_{i j}$
5) Solve matrix equation $\boldsymbol{H c}=\boldsymbol{S c} \boldsymbol{E}$
comment: In both Hückel and extended Hückel methods we assume the electrons move independently - i.e. there is no explicit electron-electron repulsion term, and so a simple Hartree product, $\Phi=\varphi_{1} \varphi_{2} \varphi_{3} \varphi_{4} \ldots \ldots$. is sufficient. It is perhaps surprising that this is even remotely successful, given the lengths we go to in ab initio theories to compute $\boldsymbol{J}$ and $\boldsymbol{K}$, and then in Cl or DFT to include correlation. The answer is that the use of parameters based on experiment means that the effects of electron-electron repulsion are implicitly included - sometimes called 'Nature's correlation'.

Example: $\mathrm{CH}_{4} \quad$ valence orbitals: $\mathbf{C 2 s}, \mathrm{C}_{2} p_{\mathrm{x}, \mathrm{y}, \mathrm{z}} \quad \mathrm{H} 1 \mathrm{~s}$



Work flow: 1) Construct $S$ using geometry and form of basis functions
2) Insert diagonal elements $\boldsymbol{H}_{i i}$ (usually tabulated) into $\boldsymbol{H}$
3) Calculate off-diagonal elements $\boldsymbol{H}_{i j}$ using W-H formula.
4) Construct secular determinant and solve for $\boldsymbol{E}$
5) Solve for $c_{n}$

Slater exponents: $\quad C 2 s=C 2 p=1.625 \mathrm{H} 1 \mathrm{~s}=1.20, \mathrm{C}-\mathrm{H}=1.09 \AA$
(this information is required to construct overlap matrix using similar procedure to that for $\mathrm{H}_{2}$ ).

1) Overlap matrix, $S$

|  | C2s | C2p ${ }_{\text {z }}$ | C2p ${ }_{\text {x }}$ | C2p ${ }_{\text {y }}$ | H1s(1) | H1s(2) | H1s(3) | H1s(4) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| C2s | 1 | 0 | 0 | 0 | 0.5183 | 0.5183 | 0.5183 | 0.5183 |
| C2p ${ }_{z}$ | 0 | 1 | 0 | 0 | 0.2819 | 0.2819 | -0.2819 | -0.2819 |
| C2p ${ }_{\text {x }}$ | 0 | 0 | 1 | 0 | -0.2819 | 0.2819 | 0.2819 | -0.2819 |
| C2p ${ }_{\text {y }}$ | 0 | 0 | 0 | 1 | 0.2819 | -0.2819 | 0.2819 | -0.2819 |
| H1s(1) | 0.5183 | 0.2819 | -0.2819 | 0.2819 | 1 | 0.1844 | 0.1844 | 0.1844 |
| H1s(2) | 0.5183 | 0.2819 | 0.2819 | -0.2819 | 0.1844 | 1 | 0.1844 | 0.1844 |
| H1s(3) | 0.5183 | -0.2819 | 0.2819 | 0.2819 | 0.1844 | 0.1844 | 1 | 0.1844 |
| H1s(4) | 0.5183 | -0.2819 | -0.2819 | -0.2819 | 0.1844 | 0.1844 | 0.1844 | 1 |

$$
H_{C_{2 s} c_{2 s}}=-19.44 \mathrm{eV}
$$

2) Parameters: $\boldsymbol{H}_{C_{2 p} C_{2 p}}=-\mathbf{1 0 . 6 7 e V}$

$$
H_{H_{15} H_{1 s}}=-13.60 \mathrm{eV}
$$

$$
H_{C_{2,} H_{1 s}}=K S_{i j} \frac{\left(H_{i i}+H_{j j}\right)}{2}=\frac{1}{2}(1.75)(0.5833)(-19.44-13.6)=-14.98 \mathrm{eV}
$$

3) $\quad H_{C_{2 p} H_{1 s}}=K S_{i j} \frac{\left(H_{i i}+H_{i j}\right)}{2}=\frac{1}{2}(1.75)( \pm 0.2819)(-10.67-13.6)= \pm 5.99 \mathrm{eV}$

$$
H_{H_{1 s} H_{1 s}}=K S_{i j} \frac{\left(H_{i i}+H_{i j}\right)}{2}=\frac{1}{2}(1.75)(0.1844)(-13.6-13.6)=-4.39 \mathrm{eV}
$$

|  | C2s | C2p ${ }_{\text {z }}$ | C2p ${ }_{\text {x }}$ | C2p ${ }_{\text {y }}$ | H1s(1) | H1s(2) | H1s(3) | H1s(4) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| C2s | -19.44 | 0 | 0 | 0 | -14.98 | -14.98 | -14.98 | -14.98 |
| C2p ${ }_{z}$ | 0 | -10.67 | 0 | 0 | -5.99 | -5.99 | 5.99 | 5.99 |
| C2p ${ }_{\text {x }}$ | 0 | 0 | -10.67 | 0 | 5.99 | -5.99 | -5.99 | 5.99 |
| C2p ${ }_{\text {y }}$ | 0 | 0 | 0 | -10.67 | -5.99 | 5.99 | -5.99 | 5.99 |
| H1s(1) | -14.98 | -5.99 | 5.99 | -5.99 | -13.6 | -4.39 | -4.39 | -4.39 |
| H1s(2) | -14.98 | -5.99 | -5.99 | 5.99 | -4.39 | -13.6 | -4.39 | -4.39 |
| H1s(3) | -14.98 | 5.99 | -5.99 | -5.99 | -4.39 | -4.39 | -13.6 | -4.39 |
| H1s(4) | -14.98 | 5.99 | 5.99 | 5.99 | -4.39 | -4.39 | -4.39 | -13.6 |

Jargon busting: 'Diagonalising the matrix':

The secular equations can be expressed in matrix form:
$\left(\begin{array}{ll}H_{a a} & H_{a b} \\ H_{a b} & H_{b b}\end{array}\right)\left(\begin{array}{ll}c_{a 1} & c_{a 2} \\ c_{b 1} & c_{a 2}\end{array}\right)=\left(\begin{array}{ll}S_{a a} & S_{a b} \\ \boldsymbol{S}_{a b} & S_{b b}\end{array}\right)\left(\begin{array}{cc}c_{a 1} & c_{a 2} \\ c_{b 1} & c_{a 2}\end{array}\right)\left(\begin{array}{cc}E_{1} & 0 \\ \mathbf{0} & E_{2}\end{array}\right)$
$\boldsymbol{H c}=\boldsymbol{S c} \boldsymbol{E} \quad$ a generalised eigenvalue problem (in a simple eigenvalue problem the overlap matrix is the identity)

When solving the Hückel and EH problems 'by hand', we would solve the secular determinant for $E$, then sub back in to the secular equations to determine the eigenfunctions (coefficients) (as in the allyl radical example above). The matrix of these coefficients, $\boldsymbol{c}$, has the property that $\boldsymbol{c}^{\dagger} \boldsymbol{H} \boldsymbol{c}$ is diagonal, with diagonal elements equal to the eigenvalues, $\boldsymbol{E}$.
Given the matrices $\boldsymbol{H}$ and $\boldsymbol{S}$, it then turns out to be easier (computationally) to determine the matrix $\boldsymbol{c}$ that makes $\boldsymbol{c}^{\dagger} \boldsymbol{H} \boldsymbol{c}$ diagonal than it is to expand the determinant itself. Thus we find the coefficients first and these give the energies, rather than vice versa. We use the expression 'diagonalising the matrix' - in effect this just means "finding the eigenvalues and eigenvectors".

## H-ES

|  | C2s | C2pz | C2p ${ }_{\text {x }}$ | C2py | H1s(1) | H1s(2) | H1s(3) | H1s(4) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| C2s | -19.44-E | 0 | 0 | 0 | -14.98-0.52E | -14.98-0.52E | -14.98-0.52E | -14.98-0.52E |
| $\mathrm{C} 2 \mathrm{p}_{\mathrm{z}}$ | 0 | -10.67-E | 0 | 0 | -5.99-0.28E | -5.99-0.28E | $5.99+0.28 \mathrm{E}$ | $5.99+0.28 \mathrm{E}$ |
| $\mathrm{C} 2 \mathrm{p}_{\mathrm{x}}$ | 0 | 0 | -10.67-E | 0 | $5.99+0.28 \mathrm{E}$ | -5.99-0.28E | $-5.99-0.28 \mathrm{E}$ | $5.99+0.28 \mathrm{E}$ |
| C2py | 0 | 0 | 0 | -10.67-E | $-5.99-0.28 \mathrm{E}$ | $5.99+0.28 \mathrm{E}$ | -5.99-0.28E | $5.99+0.28 \mathrm{E}$ |
| H1s(1) | -14.98-0.52E | -5.99-0.28E | $5.99+0.28 \mathrm{E}$ | -5.99-0.28E | -13.6-E | -4.39-0.18E | -4.39-0.18E | -4.39-0.18E |
| H1s(2) | -14.98-0.52E | -5.99-0.28E | $-5.99+0.28 \mathrm{E}$ | $5.99+0.28 \mathrm{E}$ | -4.39-0.18E | -13.6-E | -4.39-0.18E | -4.39-0.18E |
| H1s(3) | -14.98-0.52E | $5.99+0.28 \mathrm{E}$ | -5.99-0.28E | -5.99-0.28E | -4.39-0.18E | -4.39-0.18E | -13.6-E | -4.39-0.18E |
| H1s(4) | -14.98-0.52E | $5.99+0.28 \mathrm{E}$ | 5.99-0.28E | $5.99+0.28 \mathrm{E}$ | -4.39-0.18E | -4.39-0.18E | -4.39-0.18E | -13.6-E |

5) Results for $\mathrm{CH}_{4}$

|  | orbital | $\mathbf{1}$ | $\mathbf{2}$ | $\mathbf{3}$ | $\mathbf{4}$ | $\mathbf{5}$ | $\mathbf{6}$ | $\mathbf{7}$ |  |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Eigenvalues (energies) |  |  |  |  |  |  |  |  |
| $\mathbf{E}$ | -23.21 | -14.93 | -14.93 | -14.93 | 6.06 | 6.06 | 6.06 | 34.0 |  |
| Occup | 2 | 2 | 2 | 2 | 0 | 0 | 0 | 0 |  |
|  | Eigenfunctions (MOs) |  |  |  |  |  |  |  |  |
| C2s | 0.58 | 0 | 0 | 0 | 0 | 0 | 0 | 1.70 |  |
| $\mathbf{C 2 p}_{\mathbf{z}}$ | 0 | -0.12 | 0.48 | 0.20 | 0.25 | -0.31 | 1.09 | 0 |  |
| $\mathbf{C 2 p}_{\mathbf{x}}$ | 0 | -0.28 | 0.12 | -0.44 | -0.92 | 0.60 | 0.38 | 0 |  |
| C2p $_{\mathbf{y}}$ | 0 | 0.44 | 0.20 | -0.23 | -0.66 | -0.95 | -0.12 | 0 |  |
| H1s(1) | 0.19 | 0.17 | -0.10 | -0.52 | 1.0 | 0.02 | 0.45 | -0.69 |  |
| H1s(2) | 0.19 | 0.36 | 0.34 | 0.25 | -0.28 | 1.01 | 0.33 | -0.69 |  |
| H1s(3) | 0.19 | -0.50 | 0.24 | 0.0 | 0.0 | -0.67 | -0.86 | -0.69 |  |
| H1s(4) | 0.19 | -0.03 | -0.48 | 0.28 | -0.73 | 0.36 | 0.74 | -0.69 |  |
|  |  |  |  | 38 |  |  |  |  |  |



## Mulliken population analysis: where are the electrons?

Consider a normalised orbital

$$
\phi_{i}=c_{a i} \chi_{a}+c_{b i} \chi_{b} \text { and }\left\langle\phi_{i} \mid \phi_{i}\right\rangle=c_{a i}{ }^{2}+c_{b i}{ }^{2}+2 c_{a i} c_{b i} S_{a b}=1
$$

We can define net population of atom a, $\boldsymbol{q}_{a}^{i}$, and net overlap population between atoms a and $\mathrm{b}, \boldsymbol{p}_{a b}^{i}$, arising from this orbital as
$\boldsymbol{q}_{a}^{i}=c_{a i}{ }^{2} \quad \boldsymbol{p}_{a b}^{i}=\mathbf{2} c_{a i} c_{b i} \boldsymbol{S}_{a b}$
Summing these terms over all occupied orbitals leads to the atomic populations and overlap populations between centres.

Summing atomic populations for the different orbitals on a given atom ( $2 \mathrm{~s}, 2 \mathrm{p}_{\mathrm{x}, \mathrm{y}, \mathrm{z}}$ on C , for example) gives the reduced net atomic orbital and reduced overlap populations.

## In $\mathrm{CH}_{4}$ :

consider $\phi_{1}=0.58 C_{2 s}+\mathbf{0 . 1 9}\left(H_{1 s}(\mathbf{1})+H_{1 s}(2)+H_{1 s}(3)+H_{1 s}(4)\right)$ :

$$
\begin{aligned}
& q_{C}=0.58^{2}=0.34 \\
& q_{H(1)}=q_{H(2)}=q_{H(3)}=q_{H(4)}=0.19^{2}=0.04 \\
& p_{C H(1)}=p_{C H(2)}=p_{C H(3)}=p_{C H(4)}=2 \times 0.58 \times 0.19 \times 0.51=0.11 \\
& p_{H(1) H(2)}=p_{H(1) H(3)}=p_{H(1) H(4)}=p_{H(2) H(3)}=p_{H(2) H(4)}=p_{H(3) H(4)}=2 \times 0.19 \times 0.19 \times 0.18=0.01
\end{aligned}
$$

Mulliken defined gross atomic populations by dividing the overlap population equally between the two component atoms. Doing so in $\mathrm{CH}_{4}$ leads to

|  | Gross Pop | Charge |
| :--- | :--- | :--- |
| C | 3.966 | 0.034 |
| H | 1.008 | -0.008 |

Note: Mulliken populations are notoriously basis-set dependent - only good for comparisons within closely related molecules.

## Example 2 (one you can do by hand) Extended Hückel calculation on $\mathbf{N}_{2}$.

Slater exponents: $\quad \mathbf{N} 2 s=\mathbf{N} 2 p=1.95 \quad \mathbf{N}-\mathbf{N}=1.1 \AA$

1) Overlap matrix, $S$

|  | N1 2s | N1 $2 p_{\text {z }}$ | N1 $2 p_{\text {x }}$ | N1 2p ${ }^{\text {y }}$ | N2 2s | N2 2pz | N2 2p ${ }_{\text {x }}$ | N2 2p ${ }^{\text {y }}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N1 2s | 1 | 0 | 0 | 0 | 0.45 | -0.44 | 0 | 0 |
| N1 2 $p_{z}$ | 0 | 1 | 0 | 0 | 0.44 | -0.32 | 0 | 0 |
| N1 2p ${ }_{\text {x }}$ | 0 | 0 | 1 | 0 | 0 | 0 | 0.28 | 0 |
| N1 2p ${ }_{\text {y }}$ | 0 | 0 | 0 | 1 | 0 | 0 | 0 | 0.28 |
| N2 2 s | 0.45 | 0.44 | 0 | 0 | 1 | 0 | 0 | 0 |
| N2 2p $\mathrm{z}_{\mathrm{z}}$ | -0.44 | -0.32 | 0 | 0 | 0 | 1 | 0 | 0 |
| N2 2p ${ }_{\text {x }}$ | 0 | 0 | 0.28 | 0 | 0 | 0 | 1 | 0 |
| N2 2p ${ }_{\mathrm{y}}$ | 0 | 0 | 0 | 0.28 | 0 | 0 | 0 | 1 |
|  | 2) | Parameters: $\begin{aligned} H_{N_{2 s} N_{2 s}} & =-26.0 \mathrm{eV} \\ H_{N_{2 p} N_{2 p}} & =-13.4 \mathrm{eV} \end{aligned}$ | $\begin{aligned} & H_{N_{2 s} N_{2 s}}=-26.0 \mathrm{eV} \\ & H_{N_{2 p} N_{2 p}}=-13.4 \mathrm{eV} \end{aligned}$ |  |  |  |  |  |

$$
\begin{aligned}
H_{N_{2 s} N_{2 s}} & =K S_{i j} \frac{\left(H_{i i}+H_{j j}\right)}{2}=\frac{1}{2}(1.75)(0.45)(-2 * 26.0)=-20.3 \mathrm{eV} \\
H_{N_{2 p z} N_{2 p z}} & =K S_{i j} \frac{\left(H_{i i}+H_{j j}\right)}{2}=\frac{1}{2}(1.75)(-0.32)(-13.4 * 2)=7.5 \mathrm{eV} \\
3) \quad H_{N_{2 p x} N_{2 p x}} & =K S_{i j} \frac{\left(H_{i i}+H_{j j}\right)}{2}=\frac{1}{2}(1.75)(0.28)(-13.4 * 2)=-6.5 \mathrm{eV} \\
H_{N_{2 p z} N_{2 s}} & =K S_{i j} \frac{\left(H_{i i}+H_{j j}\right)}{2}=\frac{1}{2}(1.75)( \pm 0.44)(-13.4-26.0)= \pm 15.1 \mathrm{eV}
\end{aligned}
$$

|  | N1 2s | N1 ${ }^{\text {2 }} \mathrm{p}_{\mathbf{z}}$ | N1 2 $p_{\text {x }}$ | N1 2py | N2 2s | N2 2pz | N2 2p ${ }^{\text {x }}$ | N2 2py |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| N1 2s | -26.0 | 0 | 0 | 0 | -20.3 | 15.1 | 0 | 0 |
| N1 2 $p_{z}$ | 0 | -13.4 | 0 | 0 | -15.1 | 7.5 | 0 | 0 |
| N1 2p ${ }_{\text {x }}$ | 0 | 0 | -13.4 | 0 | 0 | 0 | -6.5 | 0 |
| N1 2p ${ }^{\text {y }}$ | 0 | 0 | 0 | -13.4 | 0 | 0 | 0 | -6.5 |
| N2 2s | -20.3 | -15.1 | 0 | 0 | -26.0 | 0 | 0 | 0 |
| N2 2p $\mathrm{z}_{\mathrm{z}}$ | 15.1 | 7.5 | 0 | 0 | 0 | -13.4 | 0 | 0 |
| N2 2p ${ }_{\text {x }}$ | 0 | 0 | -6.5 | 0 | 0 | 0 | -13.4 | 0 |
| N2 2p ${ }^{\text {y }}$ | 0 | 0 | 0 | -6.5 | 0 | 0 | 0 | -13.4 |

Solving (using linear algebra package of your choice) gives energy levels:

| $2 \sigma_{u}:$ | +54.0 eV |
| :--- | :--- |
| $1 \pi_{\mathrm{g}}$ | -9.5 eV |
| $2 \sigma_{\mathrm{g}}:$ | -13.5 eV |
| $1 \pi_{\mathrm{u}}:$ | -15.6 eV |
| $1 \sigma_{\mathrm{u}}:$ | -19.9 eV |
| $1 \sigma_{\mathrm{g}}:$ | -32.1 eV |$\quad$ note $2 \sigma_{\mathrm{g}}$ is above $1 \pi_{\mathrm{u}}$ (sp mixing)

## Problems sheet

## Problems 1

1.1 Write down the Slater determinant for the ground states of the following:
$\mathrm{Li}, \mathrm{He}_{2}$ and $\mathrm{Li}_{2}$.
1.2 The z component of the spin angular momentum operator is $\sum_{j=1}^{n} \hat{S}_{z j}$

Show that $\Phi=\frac{1}{\sqrt{3!}}\left|\begin{array}{lll}1 s(1) & 1 \bar{s}(\mathbf{1}) & 2 s(1) \\ 1 s(2) & 1 \bar{s}(2) & 2 s(2) \\ 1 s(3) & 1 \bar{s}(3) & 2 s(3)\end{array}\right|$ is an eigenfunction of $S_{z}$ and evaluate the eigenvalue.
1.3 By expanding the expression $\langle\Phi| \hat{\boldsymbol{H}}|\Phi\rangle=\boldsymbol{E}=\langle\Phi| \hat{\boldsymbol{H}}_{\mathbf{1}}+\hat{\boldsymbol{H}}_{2}+\frac{\boldsymbol{j}_{0}}{\boldsymbol{r}_{\mathbf{1 2}}}|\Phi\rangle$

Show that the energy of the doubly excited configuration, $\Phi=\left|1 \sigma_{u} 1 \sigma_{u}\right|$, is $E=2 E_{1 \sigma_{u}}+J_{1 \sigma_{u} 1 \sigma_{u}}$
1.4 The expression $\boldsymbol{E}=\boldsymbol{E}_{1 \sigma_{g}}+\boldsymbol{E}_{1 \sigma_{u}}+\boldsymbol{J}_{1 \sigma_{g} 1 \sigma_{u}}-\boldsymbol{K}_{1 \sigma_{g} 1 \sigma_{u}}$ was derived in the handout for the first excited triplet state of $\mathrm{H}_{2}, \Phi=\left|1 \sigma_{g} 1 \sigma_{u}\right|=\frac{1}{\sqrt{2}}\left|\begin{array}{ll}1 \sigma_{g}(\mathbf{1}) & \mathbf{1} \sigma_{u}(\mathbf{1}) \\ 1 \sigma_{g}(\mathbf{2}) & 1 \sigma_{u}(2)\end{array}\right|$.

The wavefunction above is in fact the $M_{S}=1$ component of the triplet state, also represented as $\frac{1}{\sqrt{2}}\left[1 \sigma_{g}(1) 1 \sigma_{u}(2)-1 \sigma_{g}(1) 1 \sigma_{u}(2)\right](\alpha(1) \alpha(2))$

The $M_{S}=0$ component of the same triplet state is:
${ }^{3} \Phi=\frac{1}{2}\left[1 \sigma_{g}(1) 1 \sigma_{u}(2)-1 \sigma_{g}(2) 1 \sigma_{u}(1)\right][\alpha(1) \beta(2)+\beta(1) \alpha(2)]$
while the $M_{S}=0$ component of the corresponding open-shell singlet excited state ( $\left.\left(1 \sigma_{g}\right)^{1}\left(1 \sigma_{u}\right)^{1}\right)$ is given by:
${ }^{1} \Phi=\frac{1}{2}\left[1 \sigma_{g}(1) 1 \sigma_{u}(2)+1 \sigma_{g}(2) 1 \sigma_{u}(1)\right][\alpha(1) \beta(2)-\beta(1) \alpha(2)]$

Show that these two wavefunctions can be represented as linear combinations of two Slater determinants.
${ }^{1 / 3} \Phi=\frac{1}{2}\left(\left|\begin{array}{cc}1 \sigma_{g}(1) & 1 \bar{\sigma}_{u}(1) \\ 1 \sigma_{g}(2) & 1 \bar{\sigma}_{u}(2)\end{array}\right| \pm\left|\begin{array}{ll}1 \bar{\sigma}_{g}(1) & 1 \sigma_{u}(1) \\ 1 \bar{\sigma}_{g}(2) & 1 \sigma_{u}(2)\end{array}\right|\right)$

By expanding, verify that the energy of the $M_{S}=0$ component of the triplet state is also $\boldsymbol{E}=\boldsymbol{E}_{1 \sigma_{g}}+\boldsymbol{E}_{1 \sigma_{u}}+\boldsymbol{J}_{1 \sigma_{g} 1 \sigma_{u}}-\boldsymbol{K}_{1 \sigma_{g} 1 \sigma_{u}}$.
Calculate the energy of the open-shell singlet state, ${ }^{1} \Phi$.
1.5 (a) Give expressions for the total energy of the following configuration in terms of 1- and 2-electron integrals, $E_{i}, J_{i j}$ and $K_{i j}$ :
2






11
(a)

(b)
(c)

(d)

(e)

(f)

Note that
(b) Express the total energies for the same species in terms of the Hartree Fock orbital energies, $\varepsilon_{i}, J_{i j}$ and $K_{i j}$.
(c) Show that the first ionization energy of Li is equal to $-\boldsymbol{\varepsilon}_{2 s}$.

## Problems 2

2.1 Evaluate the overlap matrix element between a normalised p-type Gaussian basis function located at the origin, $\phi_{1}$, and an s-type Gaussian located at ( $0,0, \mathrm{a}$ ), $\phi_{2}$. The exponent in both cases is $\alpha$.

$$
\phi_{1}=N_{1} z e^{-\alpha\left(x^{2}+y^{2}+z^{2}\right)} \quad \phi_{2}=N_{2} e^{-\alpha\left(x^{2}+y^{2}+(z-a)^{2}\right)}
$$

You may use the standard integral $\quad \int_{-\infty}^{+\infty} e^{-b x^{2}} d x=\sqrt{\frac{\pi}{b}}$
2.2 Use the Hückel approximation to set up and solve the secular determinant for the cyclobutadiene, $\mathrm{C}_{4} \mathrm{H}_{4}$. Now introduce non-zero overlap, $\boldsymbol{S}$, between adjacent atoms and again solve the secular equations - what impact does this have on the energy spectrum? (a typical value of $S$ is 0.2 )
2.3 Using the following data, set up the $2 \times 2$ overlap and Hamiltonian matrices for one component of the $\pi$ system ( $\pi$ and $\pi^{*}$ ) of CO (you need only consider the $p_{x}$ orbitals as the $\mathrm{p}_{\mathrm{y}}$ are equivalent by symmetry)

$$
S_{C_{2 p_{x}} o_{p_{x}}}=0.27, \quad S_{C_{2_{p_{x}}} c_{2 p_{x}}}=S_{O_{2 p_{x}} o_{p_{p_{x}}}}=1.0
$$

(all other overlap integrals are $1\left(S_{i i}\right)$ or zero $\left(S_{i j}\right)$ )

$$
\begin{aligned}
& H_{O_{2 p} o_{2 p}}=-14.8 \mathrm{eV}, \quad H_{C_{2 p} c_{2 p}}=-11.4 \mathrm{eV} \\
& H_{i j}=\frac{1}{2} * 1.75 * S_{i j}\left(H_{i i}+H_{i j}\right)
\end{aligned}
$$

Solve the $2 \times 2$ determinant and calculate the coefficients in the bonding and antibonding orbitals. Comment on your results.
2.4 Using the following data, set up the overlap and Hamiltonian matrices for $\mathrm{H}_{2} \mathrm{O}$ (bond length $=1 \AA$, angle $=90^{\circ}$ ).


$$
\begin{aligned}
& S_{O_{25} H_{1 s}(1)}=S_{O_{25} H_{1 s}(2)}=0.41, S_{O_{2 p_{z}} H_{1 s}(1)}=S_{O_{2 p_{z} H_{1 s}(2)}}=-\mathbf{0 . 2 0}, S_{O_{2 p_{x} H_{1 s}(1)}}=-S_{O_{2 p_{x}} H_{1 s}(2)}=-\mathbf{0 . 3 0} \\
& S_{H_{1 s}(1) H_{1 s}(2)}=\mathbf{0 . 1 6}
\end{aligned}
$$

(all other overlap integrals are either $1\left(S_{i i}\right)$ or zero $\left(S_{i j}\right)$ )

$$
\begin{aligned}
& H_{O_{2 s} o_{2 s}}=-32.3 \mathrm{eV} H_{O_{2 p} o_{2 p}}=-14.8 \mathrm{eV}, H_{H_{1 s} H_{1 s}}=-13.6 \mathrm{eV} \\
& H_{i j}=\frac{1}{2} K S_{i j}\left(H_{i i}+H_{i j}\right) \quad K=1.75
\end{aligned}
$$

Using an appropriate linear algebra package (octave, matlab, numpy), solve the $6 \times 6$ secular determinant to give the orbital energies and the corresponding coefficients in the LCAO expansion. Use the coefficients to sketch the orbitals.

By using appropriate normalized linear combinations of the 1 s orbitals on H , show that the secular determinant can be block diagonalised into $1 \times 1,2 \times 2$ and $3 \times 3$ components. Solve the $2 \times 2$ component by hand (expansion coefficients not required) and show that the energies of these two orbitals are unaffected by the change in basis.

## Appendix 1 (see also Valence lectures, pp18-23)

Linear variation theorem for $\mathrm{H}_{2}{ }^{+}$.

Multiply through by $\left(c_{a}{ }^{2} S_{a a}+2 c_{a} c_{b} S_{a b}+c_{b}^{2} S_{b b}\right)$ :

$$
\frac{\partial E}{\partial c_{a}}=\left[\left(2 c_{a} H_{a a}+2 c_{b} H_{a b}\right)-\frac{\left(c_{a}^{2} H_{a a}+2 c_{a} c_{b} H_{a b}+c_{b}^{2} H_{b b}\right)}{\left(c_{a}^{2} S_{a a}+2 c_{a} c_{b} S_{a b}+c_{b}{ }^{2} S_{b b}\right)}\left(2 c_{a} S_{a a}+2 c_{b} S_{a b}\right)\right]=0
$$

Note that $\frac{\left(c_{a}{ }^{2} H_{a a}+2 c_{a} c_{b} H_{a b}+c_{b}^{2} H_{b b}\right)}{\left(c_{a}{ }^{2} S_{a a}+2 c_{a} c_{b} S_{a b}+c_{b}{ }^{2} S_{b b}\right)}=\boldsymbol{E}$

$$
\begin{aligned}
& \frac{\partial \boldsymbol{E}}{\partial c_{a}}=\left[\left(2 c_{a} \boldsymbol{H}_{a a}+2 c_{b} \boldsymbol{H}_{a b}\right)-\boldsymbol{E}\left(2 c_{a} \boldsymbol{S}_{a a}+2 c_{b} \boldsymbol{S}_{a b}\right)\right] \\
& =\mathbf{2}\left[\boldsymbol{c}_{a}\left(\boldsymbol{H}_{a a}-\boldsymbol{E} \boldsymbol{S}_{a a}\right)+\boldsymbol{c}_{b}\left(\boldsymbol{H}_{a b}-\boldsymbol{E} \boldsymbol{S}_{a b}\right)\right]=\mathbf{0}
\end{aligned}
$$

Likewise

$$
\frac{\partial E}{\partial c_{b}}=\mathbf{2}\left[c_{b}\left(H_{b b}-E S_{b b}\right)+c_{a}\left(\boldsymbol{H}_{a b}-E S_{a b}\right)\right]=0
$$

$$
\left(\begin{array}{ll}
\boldsymbol{H}_{a a}-\boldsymbol{E} \boldsymbol{S}_{a a} & \boldsymbol{H}_{a b}-\boldsymbol{E} \boldsymbol{S}_{a b} \\
\boldsymbol{H}_{a b}-\boldsymbol{E} \boldsymbol{S}_{a b} & \boldsymbol{H}_{b b}-\boldsymbol{E} \boldsymbol{S}_{b b}
\end{array}\right)\binom{\boldsymbol{c}_{a}}{\boldsymbol{c}_{b}}=\mathbf{0}
$$

$$
\begin{aligned}
& \psi_{\text {trial }}=\sum_{n} c_{n} \chi_{n}=c_{a} \chi_{1 s a}+c_{b} \chi_{1 s b} \\
& \boldsymbol{E}_{\text {trial }}=\frac{\langle\boldsymbol{\psi}| \boldsymbol{H}|\boldsymbol{\psi}\rangle}{\langle\boldsymbol{\psi} \mid \boldsymbol{\psi}\rangle}=\frac{\left\langle\left(c_{a} \chi_{1 s a}+c_{b} \chi_{1 s b}\right)\right| \boldsymbol{H}\left|\left(c_{a} \chi_{1 s a}+c_{b} \chi_{1 s b}\right)\right\rangle}{\left\langle\left(c_{a} \chi_{1 s a}+c_{b} \chi_{1 s b}\right) \mid\left(c_{a} \chi_{1 s a}+c_{b} \chi_{1 s b}\right)\right\rangle} \\
& =\frac{\left\langle c_{a} \chi_{1 s a}\right| \boldsymbol{H}\left|c_{a} \chi_{1 s a}\right\rangle+\left\langle c_{a} \chi_{1 s a}\right| \boldsymbol{H}\left|c_{b} \chi_{1 s b}\right\rangle+\left\langle c_{b} \chi_{1 s b}\right| \boldsymbol{H}\left|c_{a} \chi_{1 s a}\right\rangle+\left\langle c_{b} \chi_{1 s b}\right| \boldsymbol{H}\left|c_{b} \chi_{1 s b}\right\rangle}{\left\langle\boldsymbol{c}_{a} \chi_{1 s a} \mid c_{a} \chi_{1 s a}\right\rangle+\left\langle c_{a} \chi_{1 s a} \mid c_{b} \chi_{1 s b}\right\rangle+\left\langle c_{b} \chi_{1 s b} \mid c_{a} \chi_{1 s a}\right\rangle+\left\langle c_{b} \chi_{1 s b} \mid c_{b} \chi_{1 s b}\right\rangle} \\
& =\frac{c_{a}^{2} H_{a a}+2 c_{a} c_{b} H_{a b}+c_{b}^{2} H_{b b}}{c_{a}{ }^{2} S_{a a}+2 c_{a} c_{b} S_{a b}+c_{b}{ }^{2} S_{b b}} \quad H_{a a}=\left\langle\chi_{1 s a}\right| H\left|\chi_{1 s a}\right\rangle, S_{a a}=\left\langle\chi_{1 s a} \mid \chi_{1 s a}\right\rangle \text { etc } \\
& \frac{\partial E}{\partial c_{a}}=\frac{\left[\left(c_{a}^{2} S_{a a}+2 c_{a} c_{b} S_{a b}+c_{b}^{2} S_{b b}\right)\left(2 c_{a} H_{a a}+2 c_{b} H_{a b}\right)-\left(c_{a}^{2} H_{a a}+2 c_{a} c_{b} H_{a b}+c_{b}^{2} H_{b b}\right)\left(2 c_{a} S_{a a}+2 c_{b} S_{a b}\right)\right]}{\left(c_{a}{ }^{2} S_{a a}+2 c_{a} c_{b} S_{a b}+c_{b}{ }^{2} S_{b b}\right)^{2}}=0 \\
& \text { (using } \frac{\partial\left(\frac{f_{1}(x, y)}{f_{2}(x, y)}\right)}{\partial x}=\frac{\left[f_{2}(x, y) f_{1}^{\prime}(x, y)-f_{1}(x, y) f_{2}^{\prime}(x, y)\right]}{f_{2}(x, y)^{2}} \text { ) }
\end{aligned}
$$

This system of simultaneous equations has non-trivial solutions when:

$$
\left|\begin{array}{ll}
\boldsymbol{H}_{a a}-\boldsymbol{E} \boldsymbol{S}_{a a} & \boldsymbol{H}_{a b}-\boldsymbol{E} \boldsymbol{S}_{a b} \\
\boldsymbol{H}_{a b}-\boldsymbol{E} \boldsymbol{S}_{a b} & \boldsymbol{H}_{b b}-\boldsymbol{E} \boldsymbol{S}_{b b}
\end{array}\right|=\mathbf{0}
$$

Generalising to arbitrary size: variationally optimised solution occurs when:

$$
\left|\boldsymbol{H}_{i j}-E S_{i j}\right|=0
$$

If we assume normalised atomic orbitals:

$$
\begin{aligned}
& \boldsymbol{S}_{a a}=\boldsymbol{S}_{b b}=\mathbf{1}, \boldsymbol{S}_{a b}=\boldsymbol{S} \\
& \boldsymbol{H}_{a a}=\boldsymbol{H}_{b b}=\boldsymbol{\alpha}, \boldsymbol{H}_{a b}=\boldsymbol{H}_{b a}=\boldsymbol{\beta}
\end{aligned}
$$

and define

$$
\begin{aligned}
& \left|\begin{array}{cc}
\boldsymbol{H}_{a a}-\boldsymbol{E} \boldsymbol{S}_{a a} & \boldsymbol{H}_{a b}-\boldsymbol{E} \boldsymbol{S}_{a b} \\
\boldsymbol{H}_{a b}-\boldsymbol{E} \boldsymbol{S}_{a b} & \boldsymbol{H}_{b b}-\boldsymbol{E} \boldsymbol{S}_{b b}
\end{array}\right|=\left|\begin{array}{cc}
\boldsymbol{\alpha}-\boldsymbol{E} & \boldsymbol{\beta}-\boldsymbol{E} \boldsymbol{S} \\
\boldsymbol{\beta}-\boldsymbol{E} & \boldsymbol{\alpha}-\boldsymbol{E}
\end{array}\right|=\mathbf{0} \\
& (\boldsymbol{\alpha}-\boldsymbol{E})^{2}-(\boldsymbol{\beta}-\boldsymbol{E} \boldsymbol{S})^{2}=\mathbf{0} \\
& \Rightarrow(\boldsymbol{\alpha}-\boldsymbol{E}+\boldsymbol{\beta}-\boldsymbol{E} \boldsymbol{S})(\boldsymbol{\alpha}-\boldsymbol{E}-\boldsymbol{\beta}+\boldsymbol{E S})=\mathbf{0} \\
& \Rightarrow \boldsymbol{E}(-\mathbf{1}-\boldsymbol{S})+(\boldsymbol{\alpha}+\boldsymbol{\beta})=\mathbf{0}
\end{aligned}
$$


(difference of 2 squares)
or $\boldsymbol{E}(-\mathbf{1}+\boldsymbol{S})+(\boldsymbol{\alpha}-\boldsymbol{\beta})=\mathbf{0}$

$$
E_{ \pm}=\frac{(\alpha \pm \beta)}{(1 \pm \boldsymbol{S})} \text { (note } \beta \text { is negative) }
$$

These are the energies of the $1 \sigma_{g}$ and $1 \sigma_{u}$ orbitals of $\mathrm{H}_{2}{ }^{+}$.
Sub $\boldsymbol{E}_{+}$back into secular equations to get $\boldsymbol{c}_{\boldsymbol{a}}, \boldsymbol{c}_{\boldsymbol{b}}$ :

$$
E_{+}=\frac{(\alpha+\beta)}{(1+S)} \Rightarrow c_{a}\left(\alpha-\left(\frac{\alpha+\beta}{1+S}\right)\right)+c_{b}\left(\beta-\left(\frac{\alpha+\beta}{1+S}\right) S\right)=0
$$

multiply through by $1+S$ :

$$
\begin{aligned}
& c_{a}((1+S) \alpha-(\alpha+\beta))+c_{b}((1+S) \beta-(\alpha+\beta) S)=0 \\
& \Rightarrow c_{a}(S \alpha-\beta)+c_{b}(\beta-S \alpha)=0 \\
& \Rightarrow c_{a}(S \alpha-\beta)=c_{b}(S \alpha-\beta) \\
& \Rightarrow c_{a}=c_{b}
\end{aligned}
$$

Likewise $\quad E_{-}=\frac{(\boldsymbol{\alpha}-\boldsymbol{\beta})}{(\mathbf{1}-\boldsymbol{S})} \Rightarrow \boldsymbol{c}_{a}=-\boldsymbol{c}_{b}$

Normalising for $\psi_{+}=\boldsymbol{c}_{a}\left(\chi_{1 s a}+\chi_{1 s b}\right)=\mathbf{1} \sigma_{g}$ :

$$
\begin{aligned}
& \int c_{a}\left(\chi_{1 s a}+\chi_{1 s b}\right) c_{a}\left(\chi_{1 s a}+\chi_{1 s b}\right) d \tau=c_{a}^{2}(1+1+S+S)=1 \\
& \Rightarrow c_{a}=c_{b}=\frac{1}{\sqrt{2(1+S)}}
\end{aligned}
$$

Likewise for $\psi_{-}=c_{a}\left(\chi_{1 s a}-\chi_{1 s b}\right)=1 \sigma_{u}$

$$
\Rightarrow c_{a}=-c_{b}=\frac{1}{\sqrt{2(1-S)}}
$$

